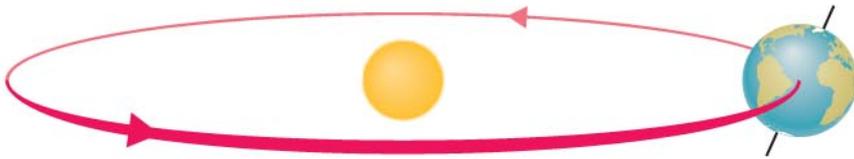
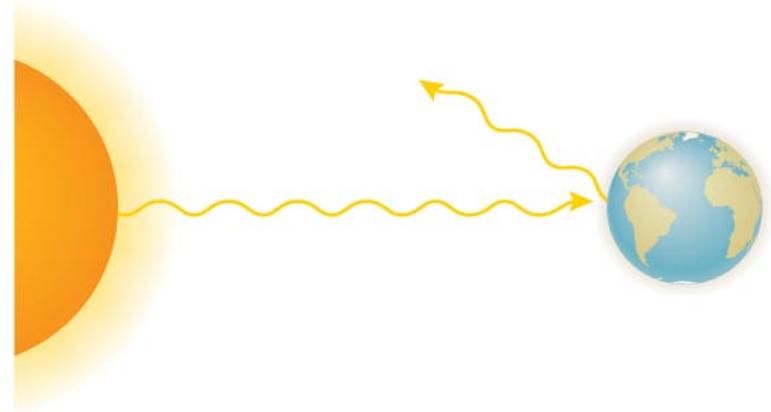
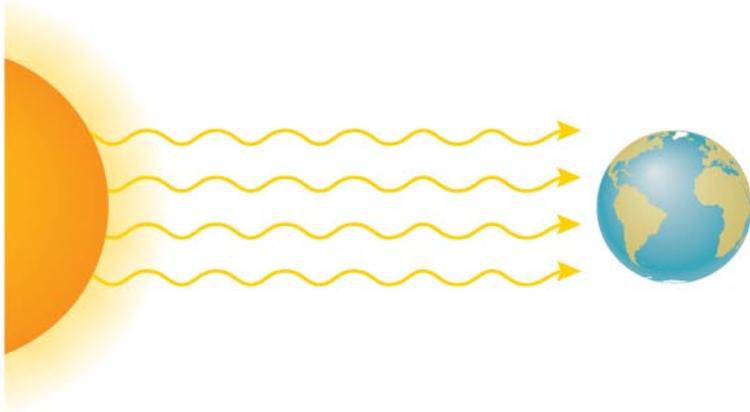
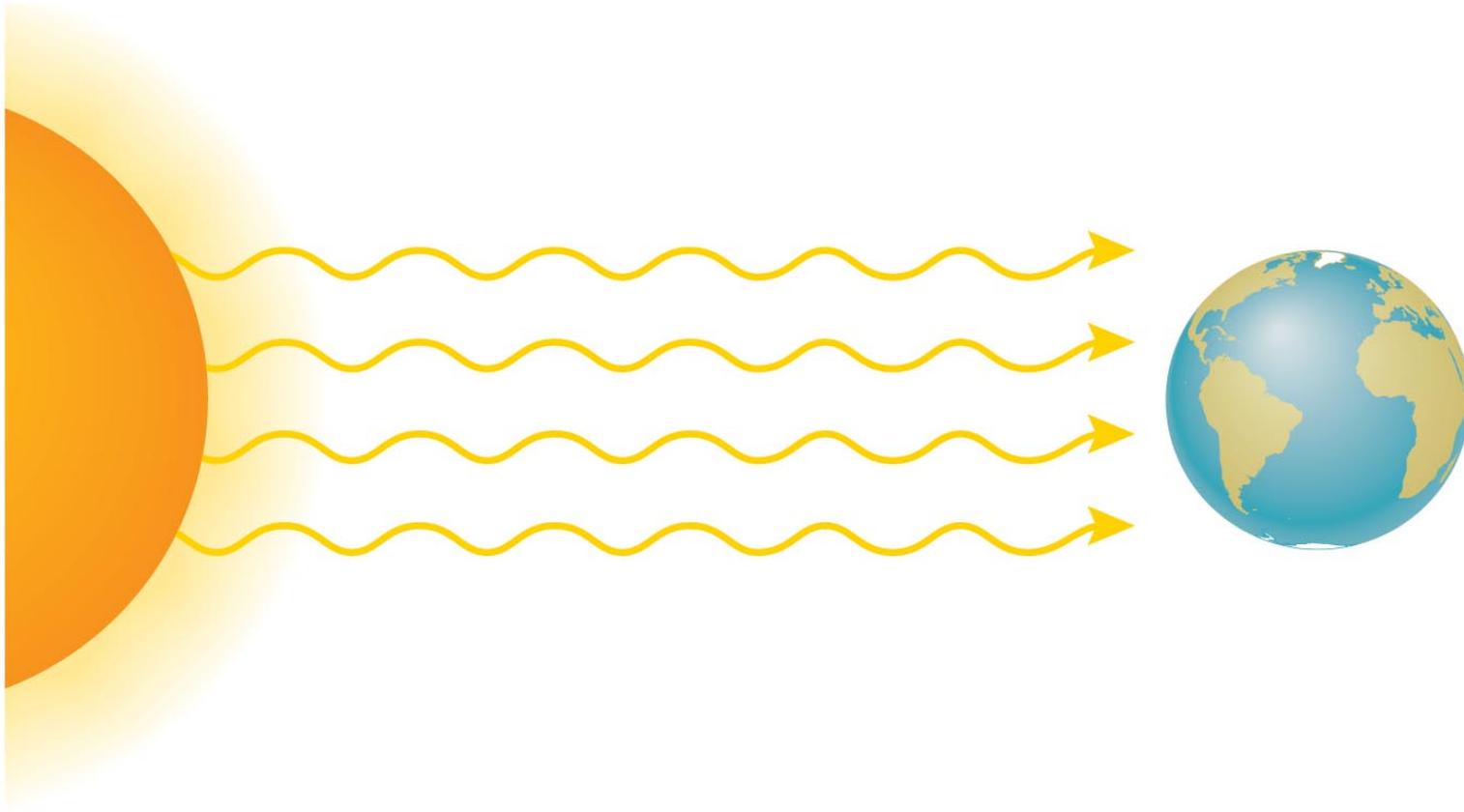


# What factors can cause long-term climate change?

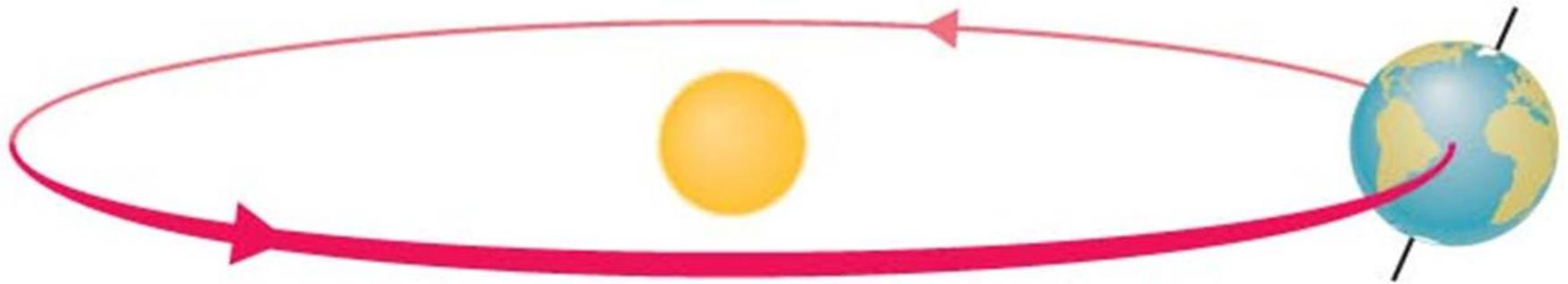


# Solar Brightening



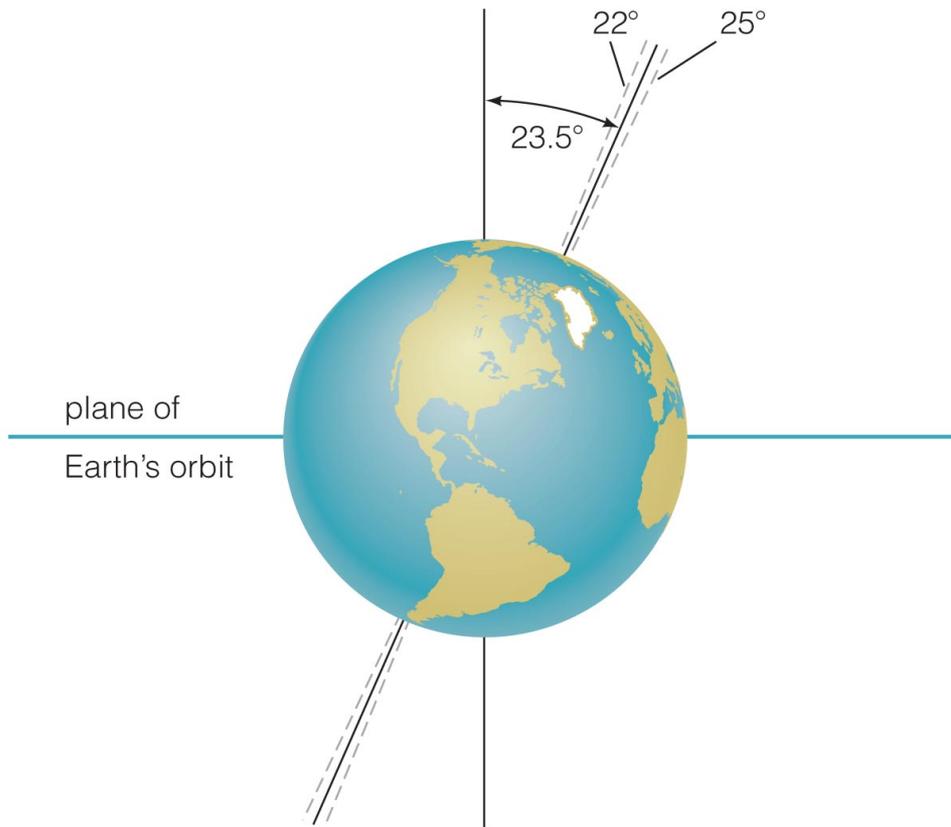
- The Sun very gradually grows brighter with time, increasing the amount of sunlight warming the planets.

# Changes in Axis Tilt



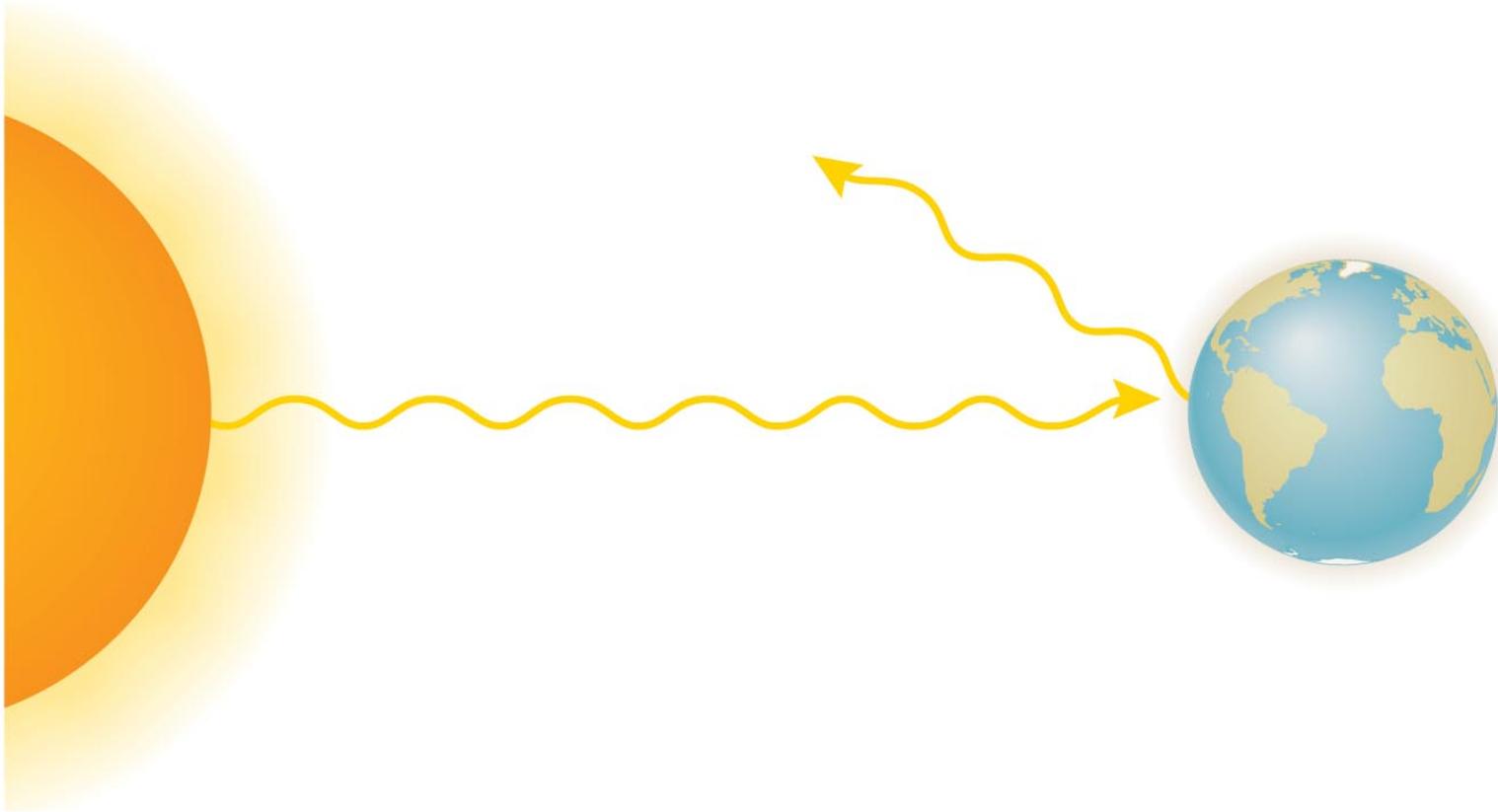
- Greater tilt creates more extreme seasons, while smaller tilt keeps polar regions colder.

# Changes in Axis Tilt



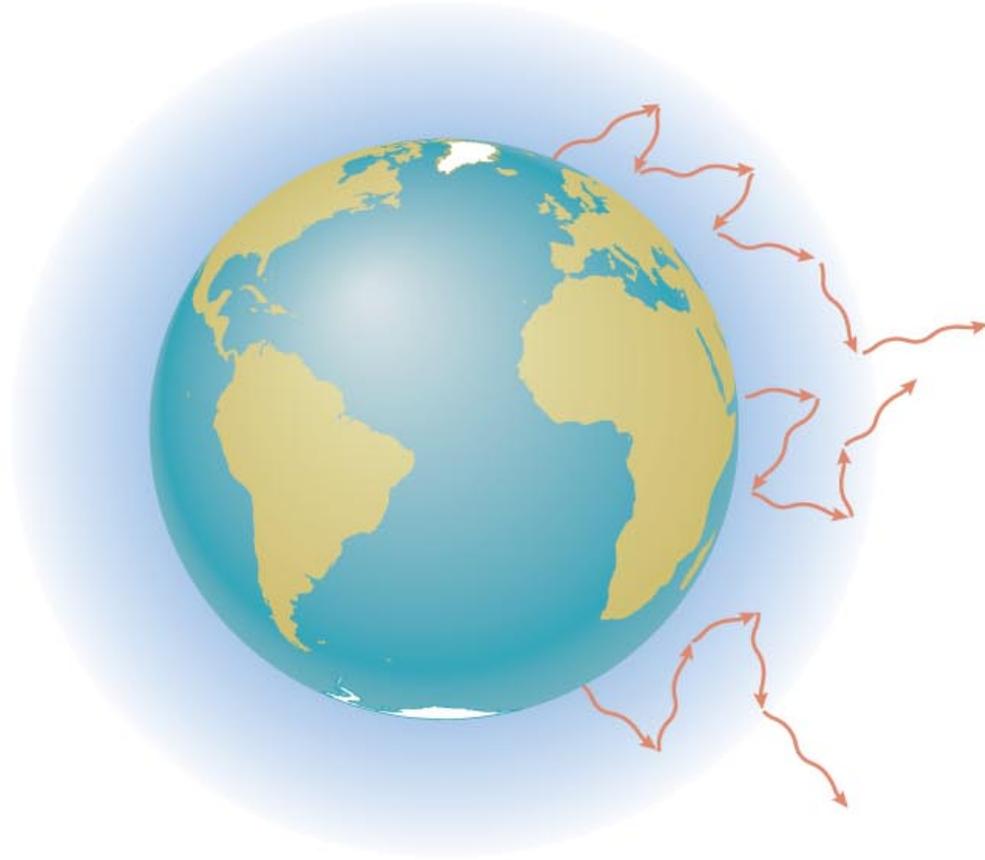
- Small gravitational tugs from other bodies in solar system cause Earth's axis tilt to vary between  $22^\circ$  and  $25^\circ$ .

# Changes in Reflectivity



- Higher reflectivity tends to cool a planet, while lower reflectivity leads to warming.

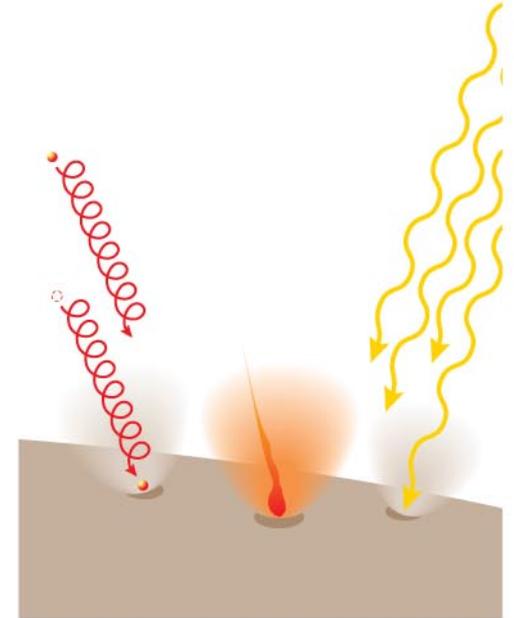
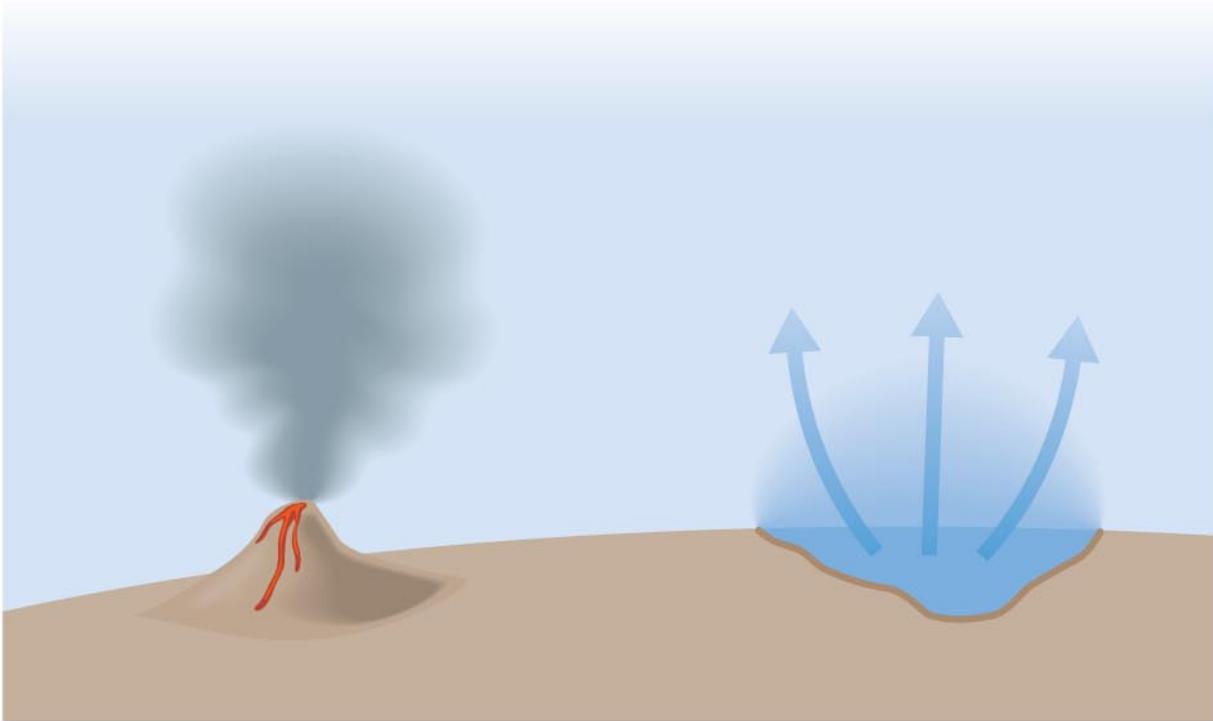
# Changes in Greenhouse Gases



- An increase in greenhouse gases leads to warming, while a decrease leads to cooling.

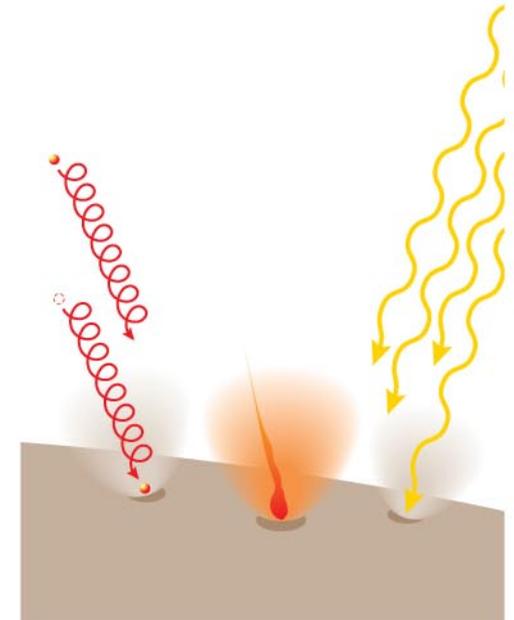
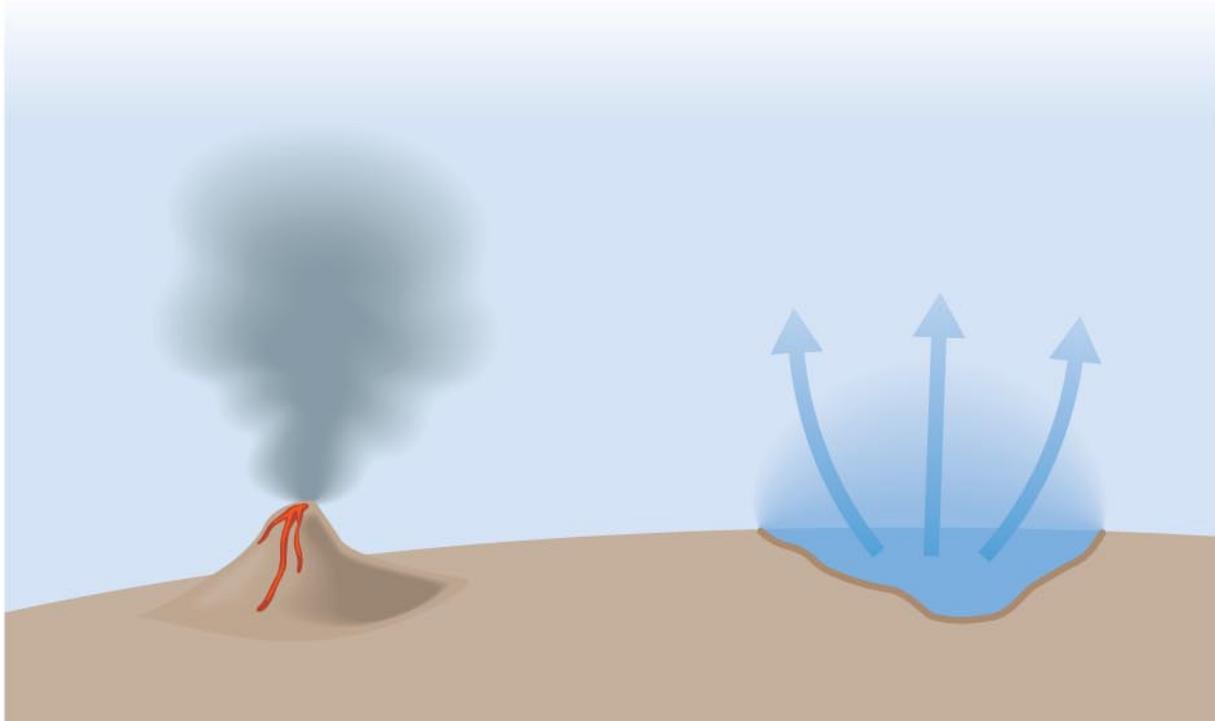
# How does a planet gain or lose atmospheric gases?

How Atmospheres Gain Gas



# Sources of Gas

## How Atmospheres Gain Gas



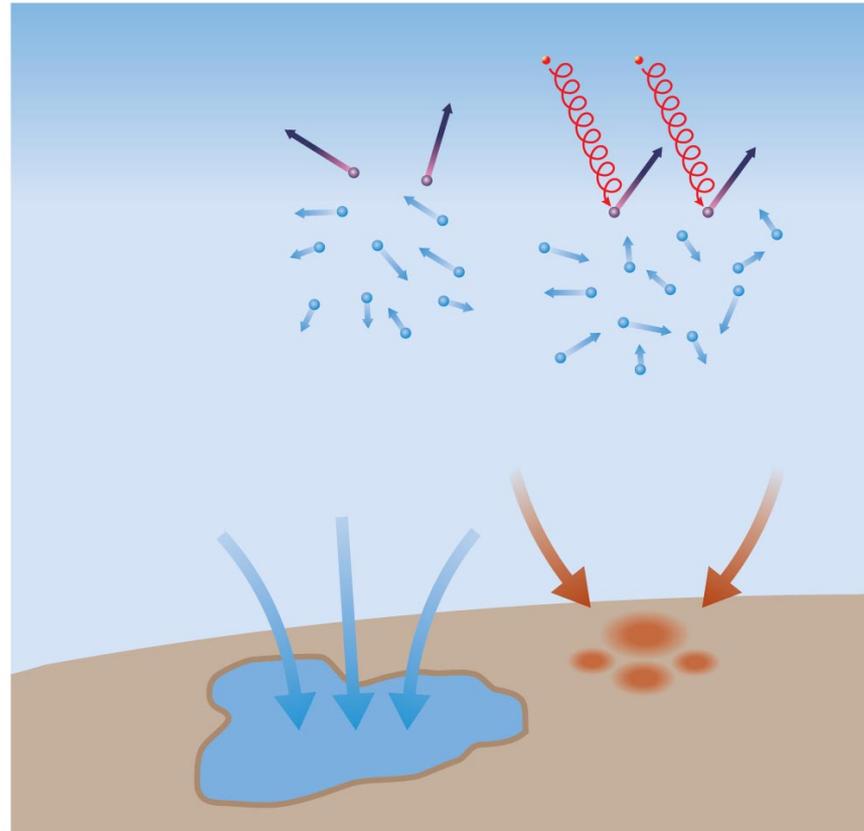
**Outgassing**  
from  
volcanoes

**Evaporation** of  
surface liquid;  
**sublimation** of  
surface ice

**Impacts** of  
particles and  
photons

# Losses of Gas

How Atmospheres Lose Gas



**Thermal escape** of atoms

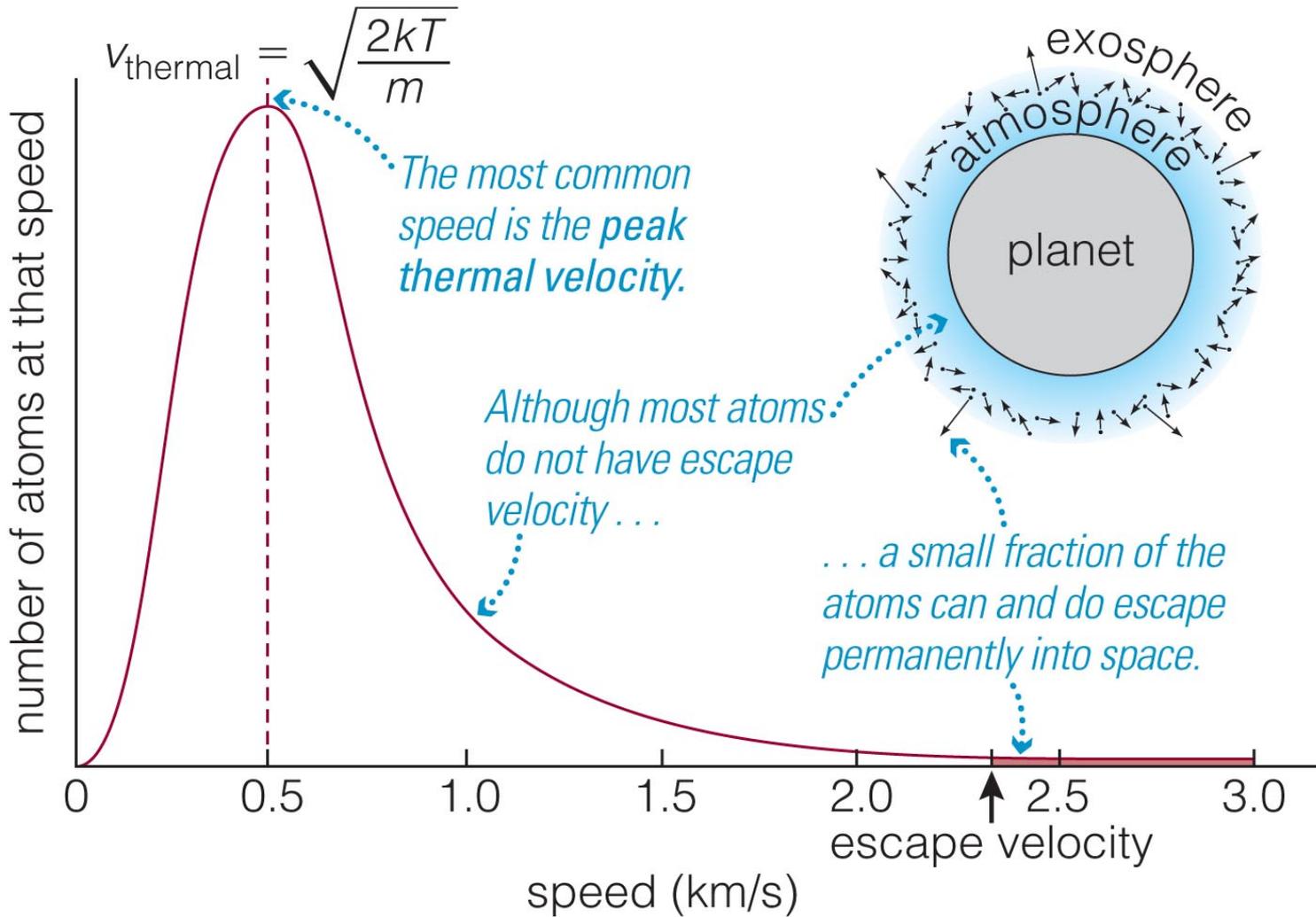
**Sweeping** by solar wind

**Condensation** onto surface

**Chemical reactions** with surface

**Large impacts** blasting gas into space

# Thermal Escape



Interactive Figure 

# What have we learned?

- **What creates wind and weather?**
  - Atmospheric heating and the Coriolis effect
- **What factors can cause long-term climate change?**
  - Brightening of the Sun
  - Changes in axis tilt
  - Changes in reflectivity
  - Changes in greenhouse gases

# What have we learned?

- **How does a planet gain or lose atmospheric gases?**
  - Gains: outgassing, evaporation/sublimation, and impacts by particles and photons
  - Losses: condensation, chemical reactions, blasting by large impacts, sweeping by solar winds, and thermal escape

# 10.3 Atmospheres of Moon and Mercury

- Our goals for learning:
  - **Do the Moon and Mercury have any atmosphere?**

# Do the Moon and Mercury have any atmosphere?

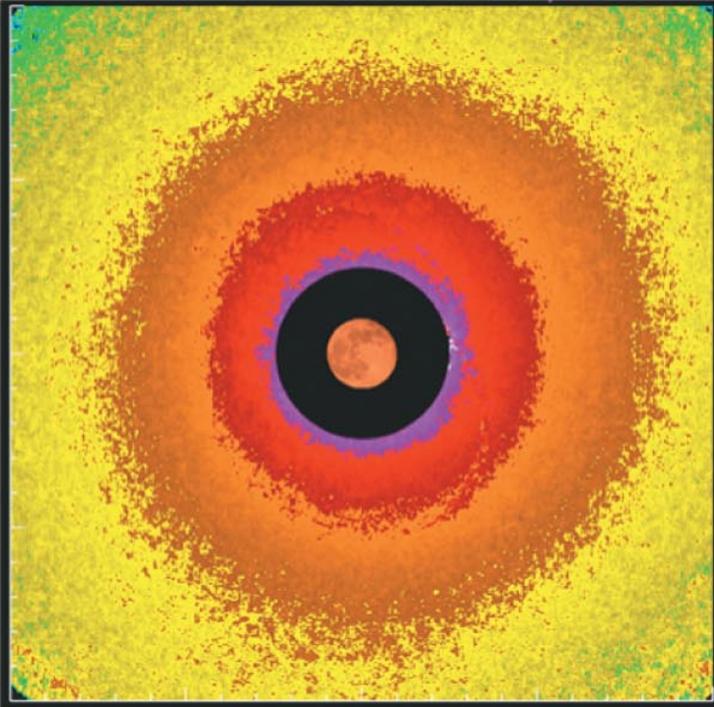


Moon

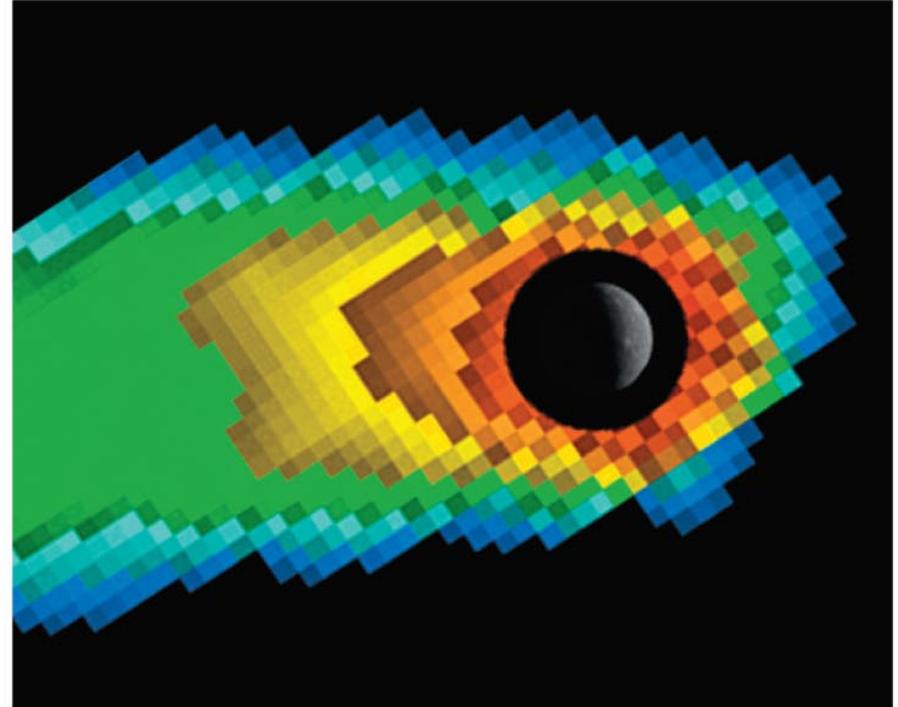


Mercury

# Exospheres of the Moon and Mercury



**a** The Moon's exosphere, which extends high above the surface.



**b** Mercury's exosphere, much of which is escaping in this image.

- Sensitive measurements show that the Moon and Mercury have extremely thin atmospheres.
- Gas comes from impacts that eject surface atoms.

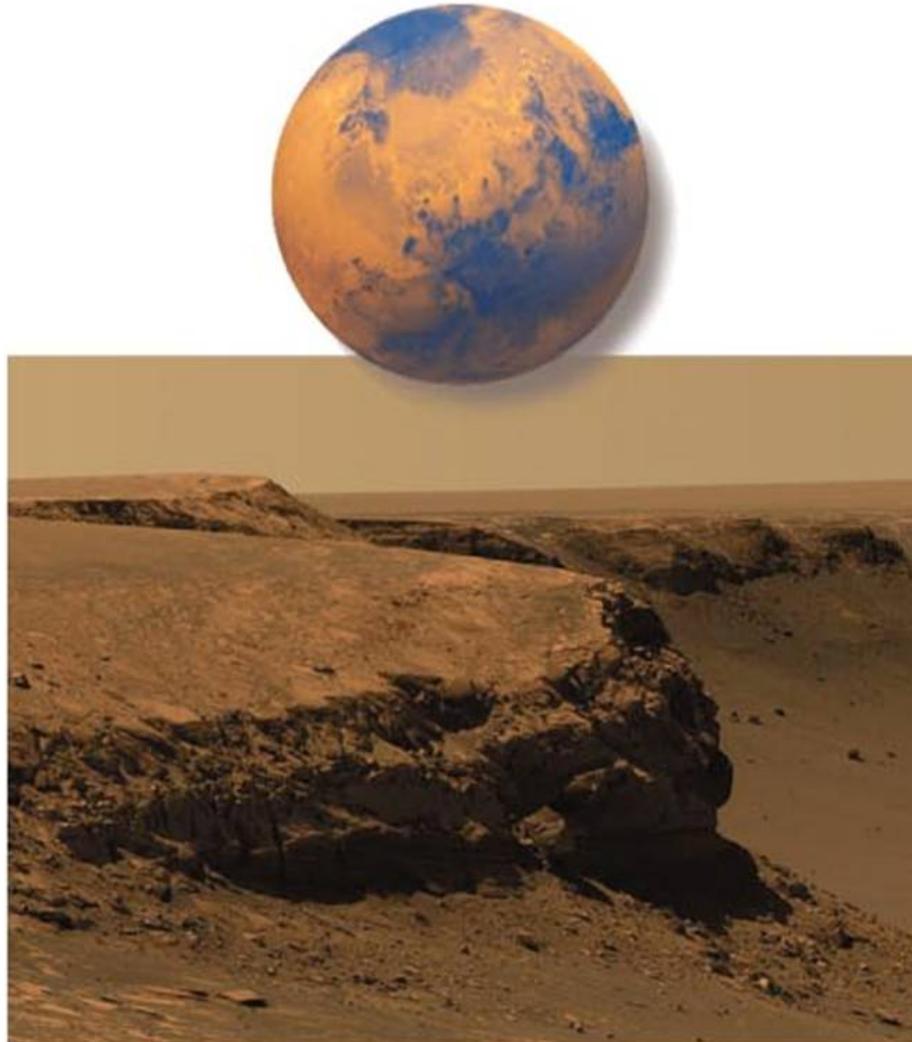
# What have we learned?

- **Do the Moon and Mercury have any atmosphere?**
  - The Moon and Mercury have very thin atmospheres made up of particles ejected from surface.

# 10.4 The Atmospheric History of Mars

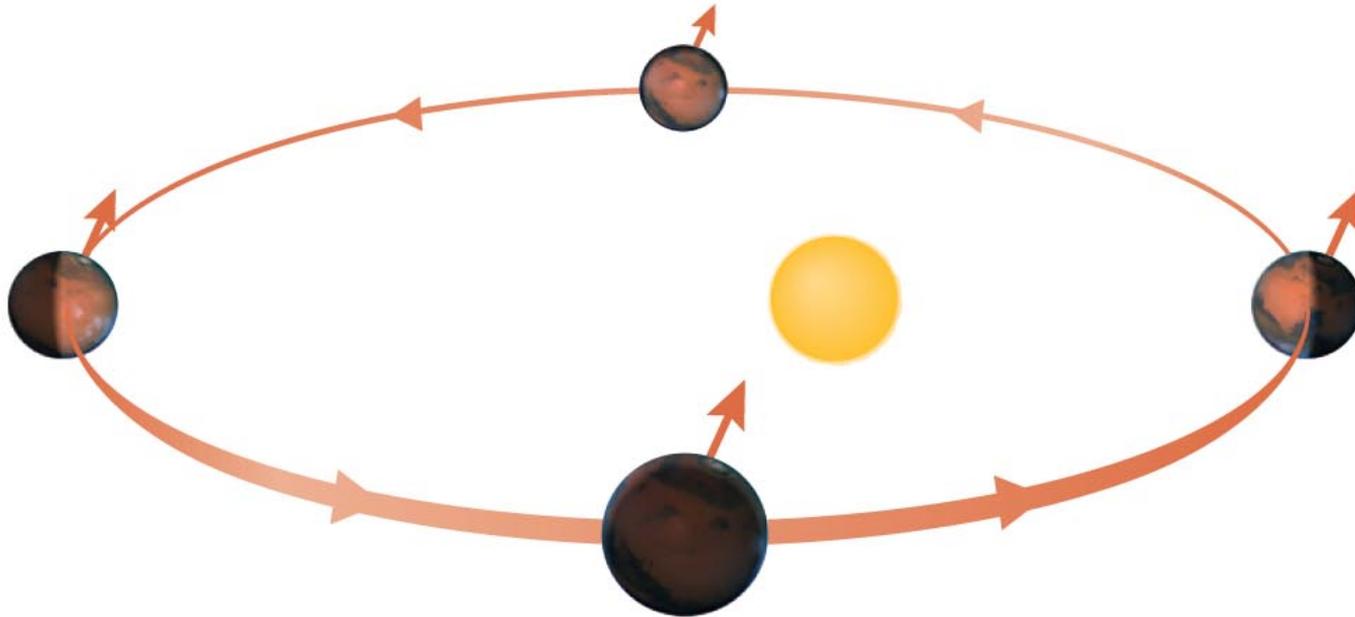
- Our goals for learning:
  - **What is Mars like today?**
  - **Why did Mars change?**

# What is Mars like today?



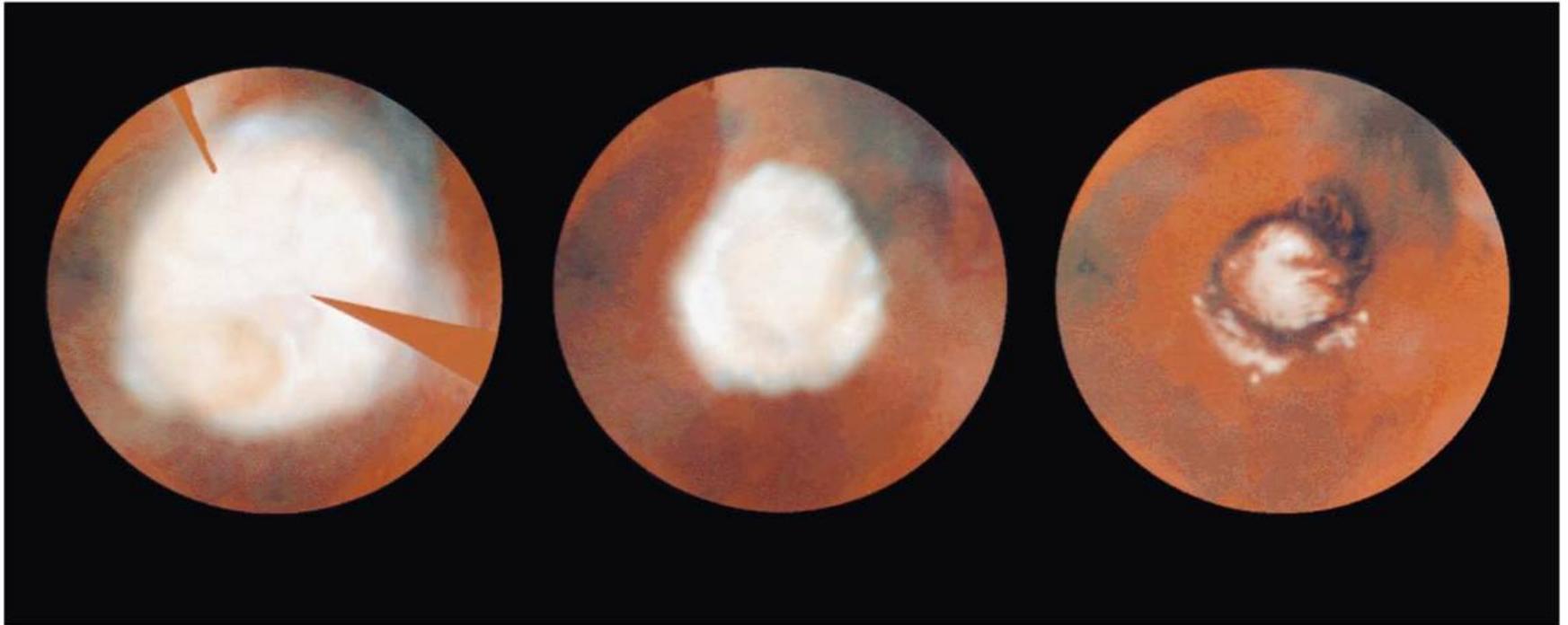
# Seasons on Mars

## Seasons on Mars



- The ellipticity of Mars's orbit makes seasons more extreme in the southern hemisphere.

# Polar Ice Caps of Mars



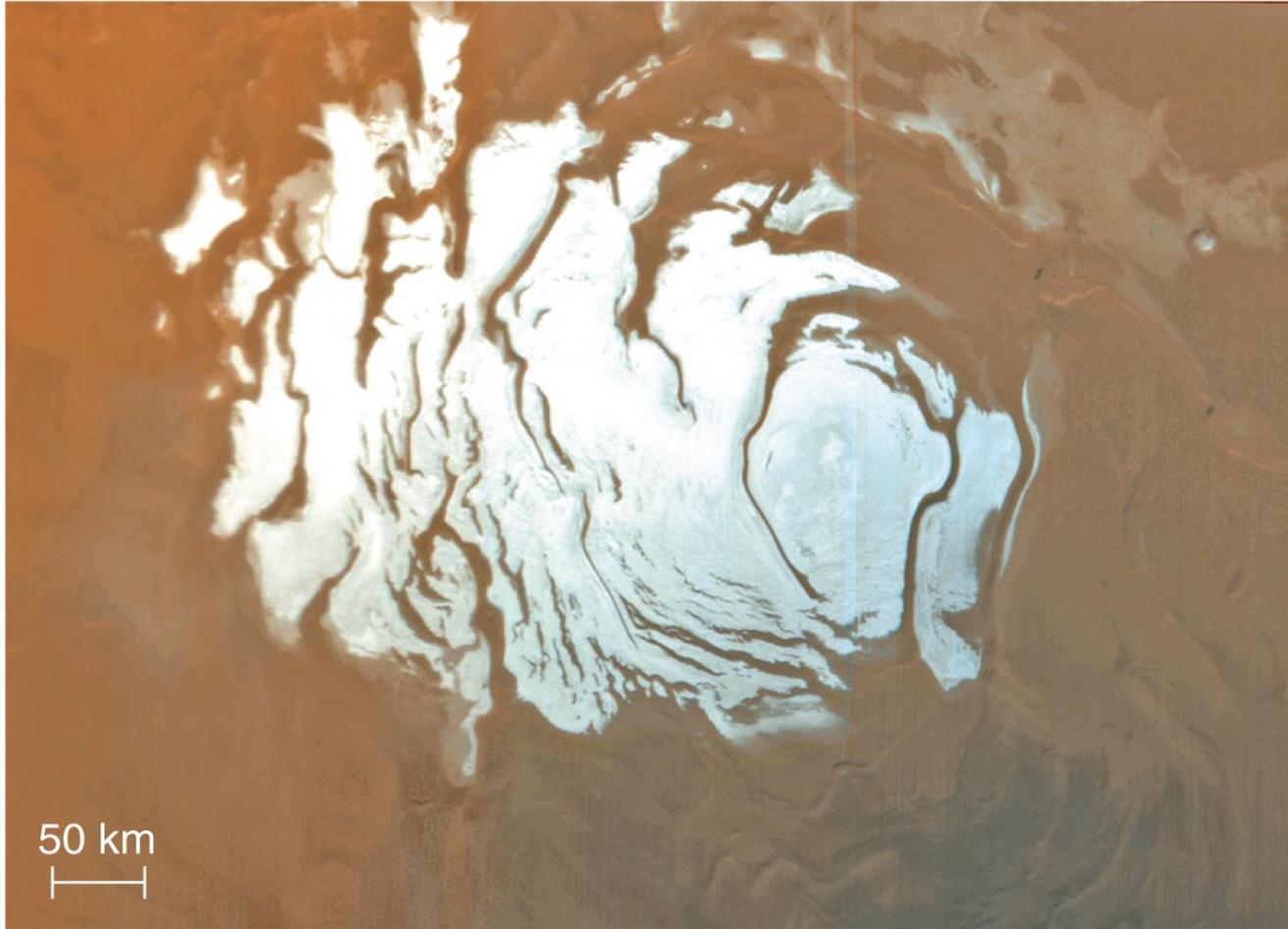
Late winter

Mid-spring

Early summer

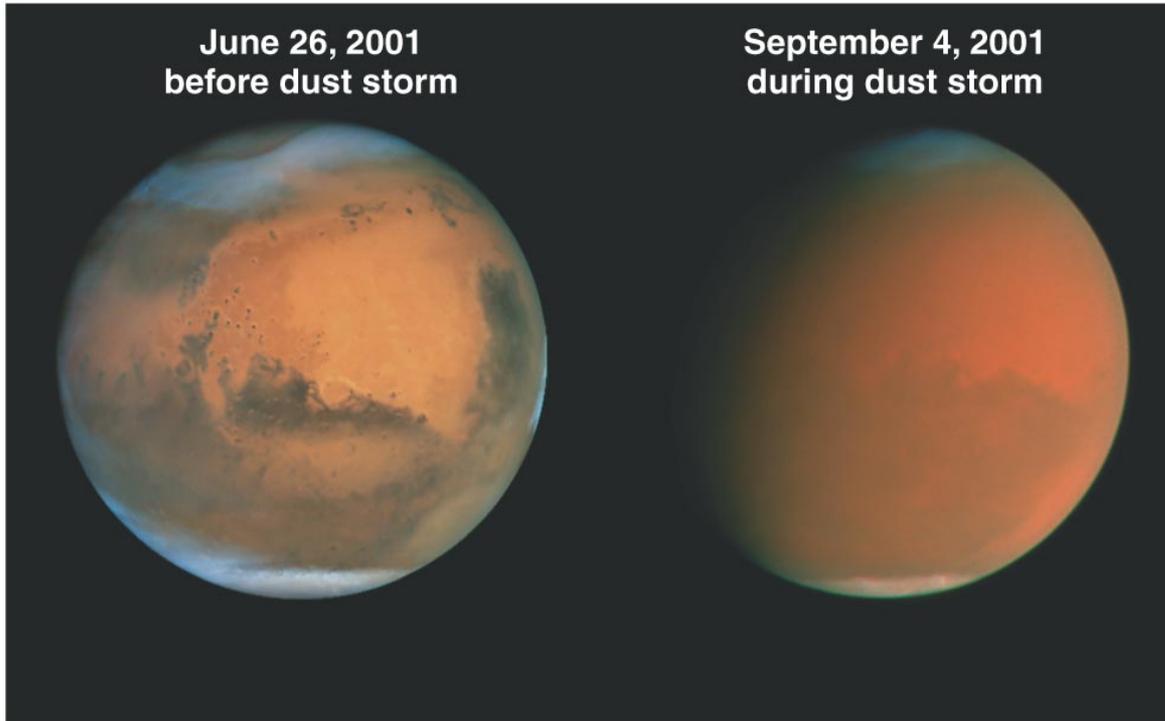
Carbon dioxide ice of polar cap sublimates as summer approaches and condenses at opposite pole.

# Polar Ice Caps of Mars



Residual ice of the south polar cap remaining during summer is primarily water ice.

# Dust Storms on Mars

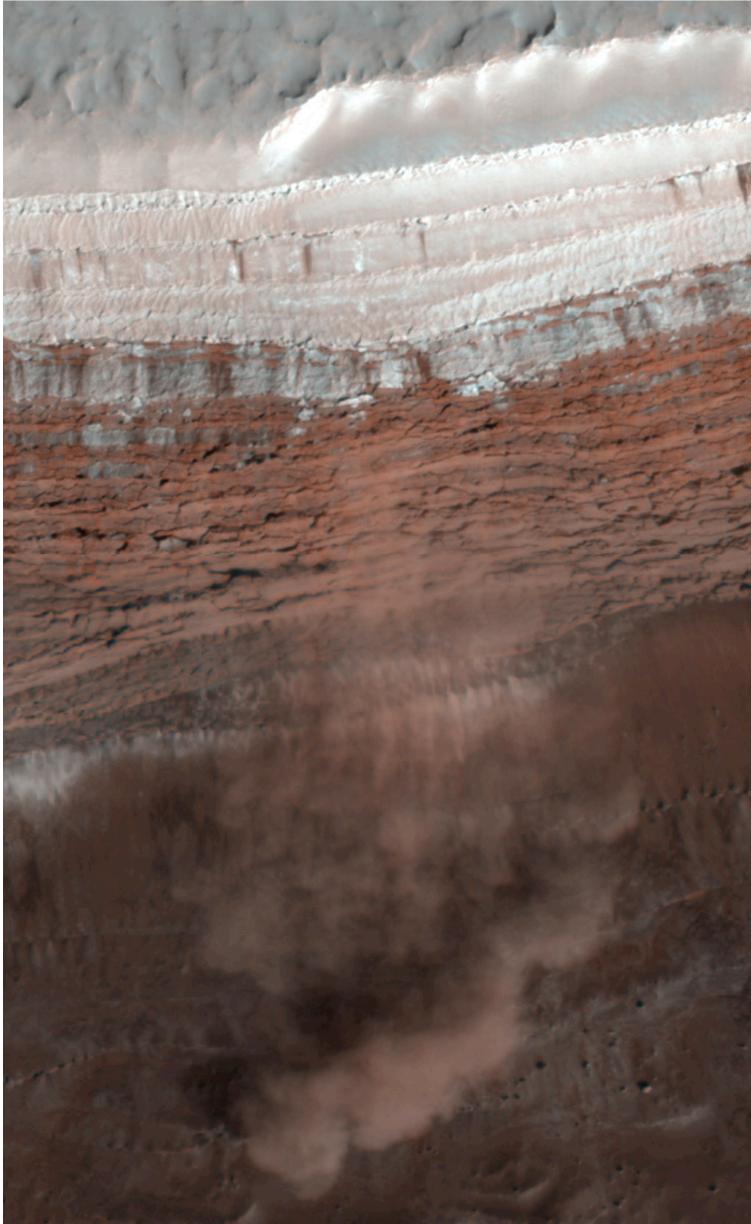


Interactive Figure

## *Martian Dust Devil*

- Seasonal winds can drive dust storms on Mars.
- Dust in the atmosphere absorbs blue light, sometimes making the sky look brownish-pink.

# Changing Axis Tilt

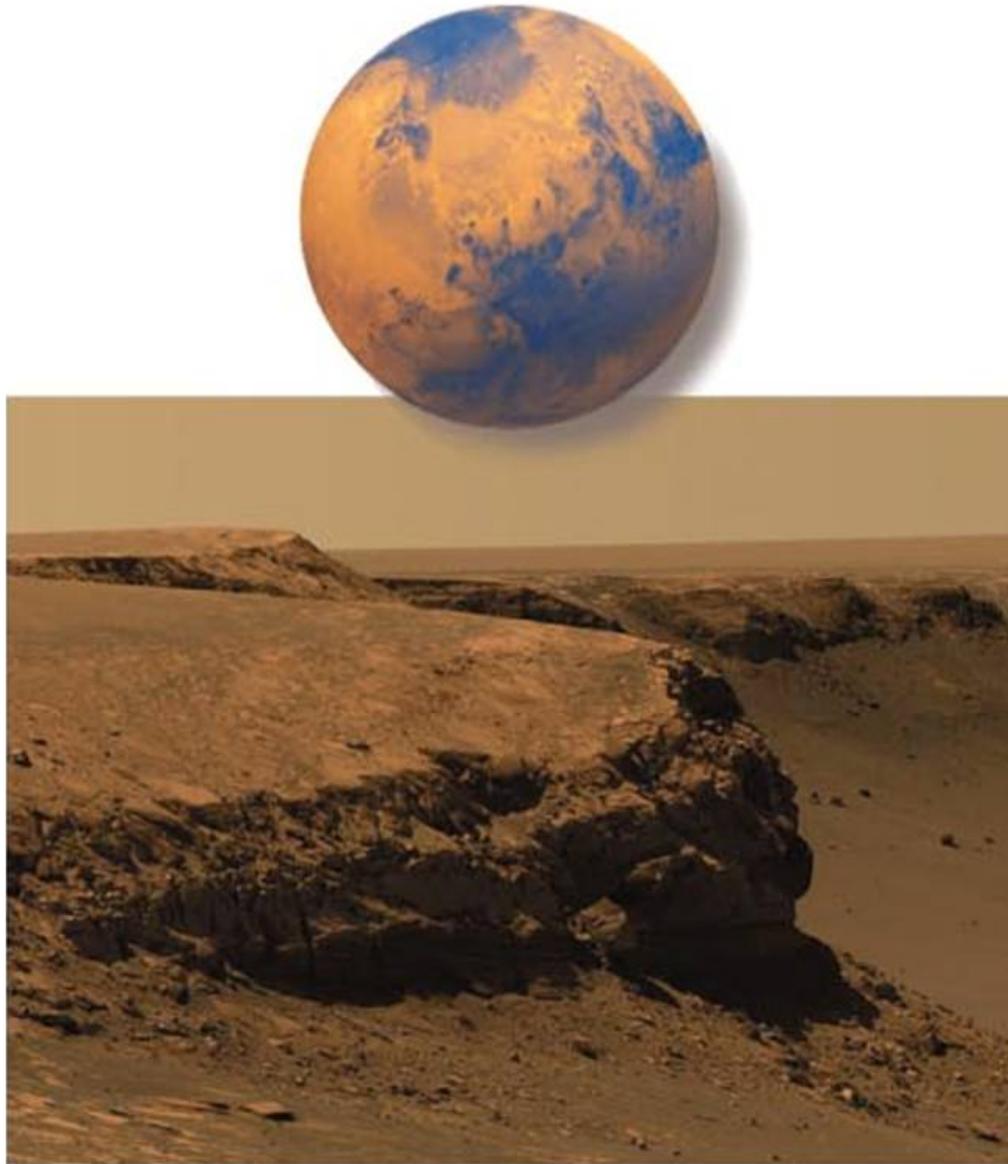


Calculations suggest Mars's axis tilt ranges from  $0^{\circ}$  to  $60^{\circ}$ .

Such extreme variations can cause climate changes.

Alternating layers of ice and dust in polar regions reflect these climate changes.

# Why did Mars change?



# Climate Change on Mars

- Mars has not had widespread surface water for 3 billion years.
- Greenhouse effect probably kept the surface warmer before that.
- Somehow Mars lost most of its atmosphere.
- Magnetic field may have preserved early Martian atmosphere.
- Solar wind may have stripped atmosphere after magnetic field decreased because of interior cooling.
- No more volcanoes, no more outgassing, no more atmosphere

# What have we learned?

- **What is Mars like today?**
  - Mars is cold, dry, and frozen.
  - Strong seasonal changes cause CO<sub>2</sub> to move from pole to pole, leading to dust storms.
- **Why did Mars change?**
  - Its atmosphere must have once been much thicker for its greenhouse effect to allow liquid water on the surface.
  - Somehow Mars lost most of its atmosphere, perhaps because of its declining magnetic field.

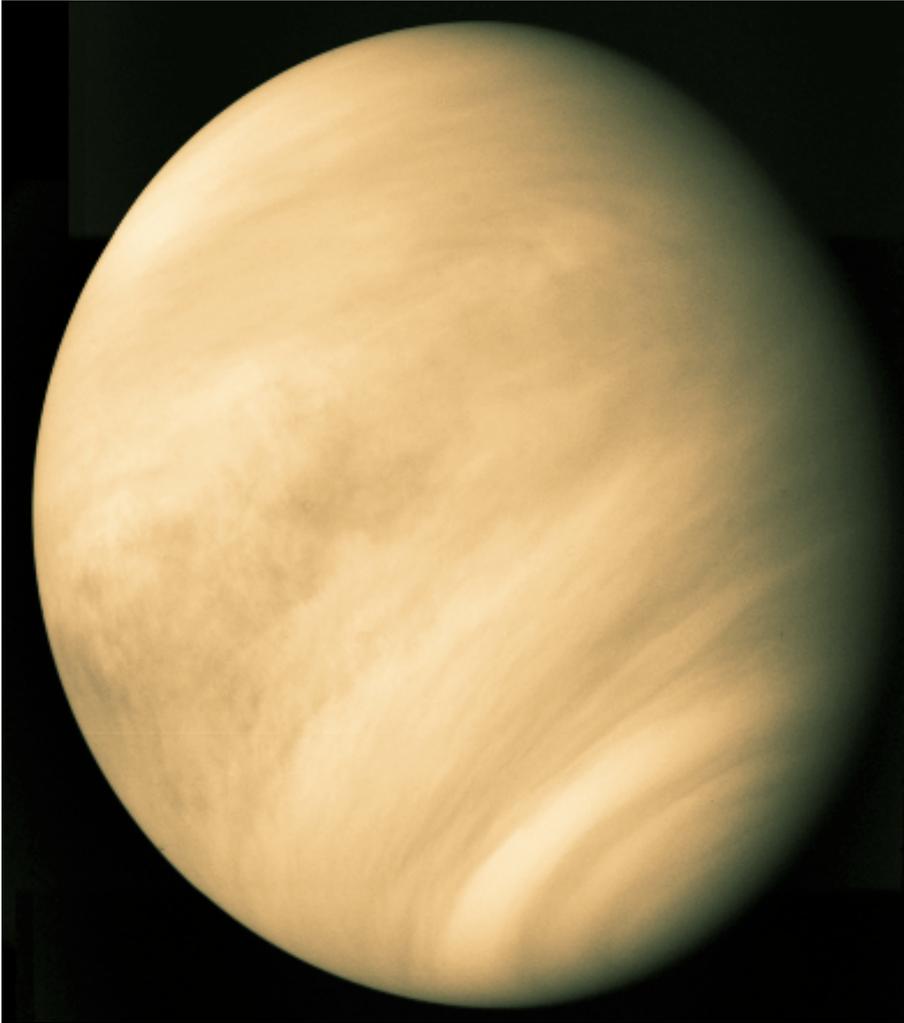
# 10.5 The Atmospheric History of Venus

- Our goals for learning:
  - **What is Venus like today?**
  - **How did Venus get so hot?**

# What is Venus like today?



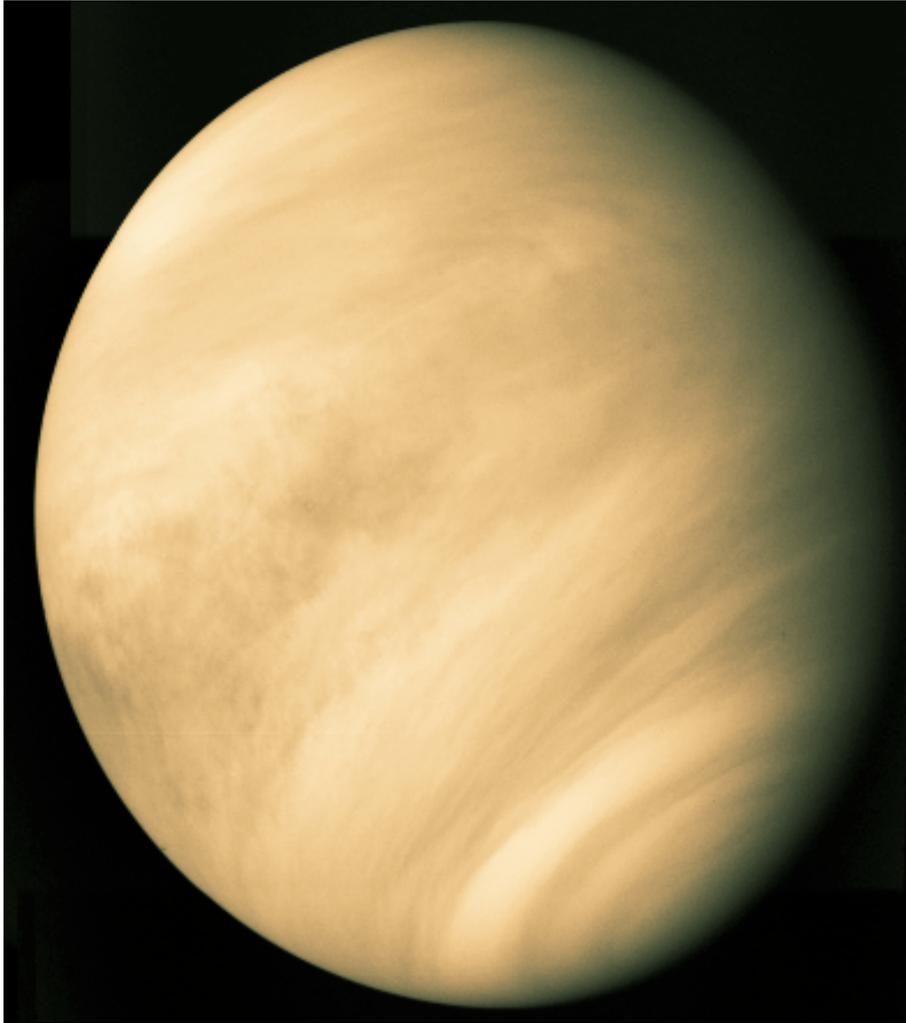
# Atmosphere of Venus



Venus has a very thick carbon dioxide atmosphere with a surface pressure 90 times that of Earth.

Slow rotation produces a very weak Coriolis effect and little weather.

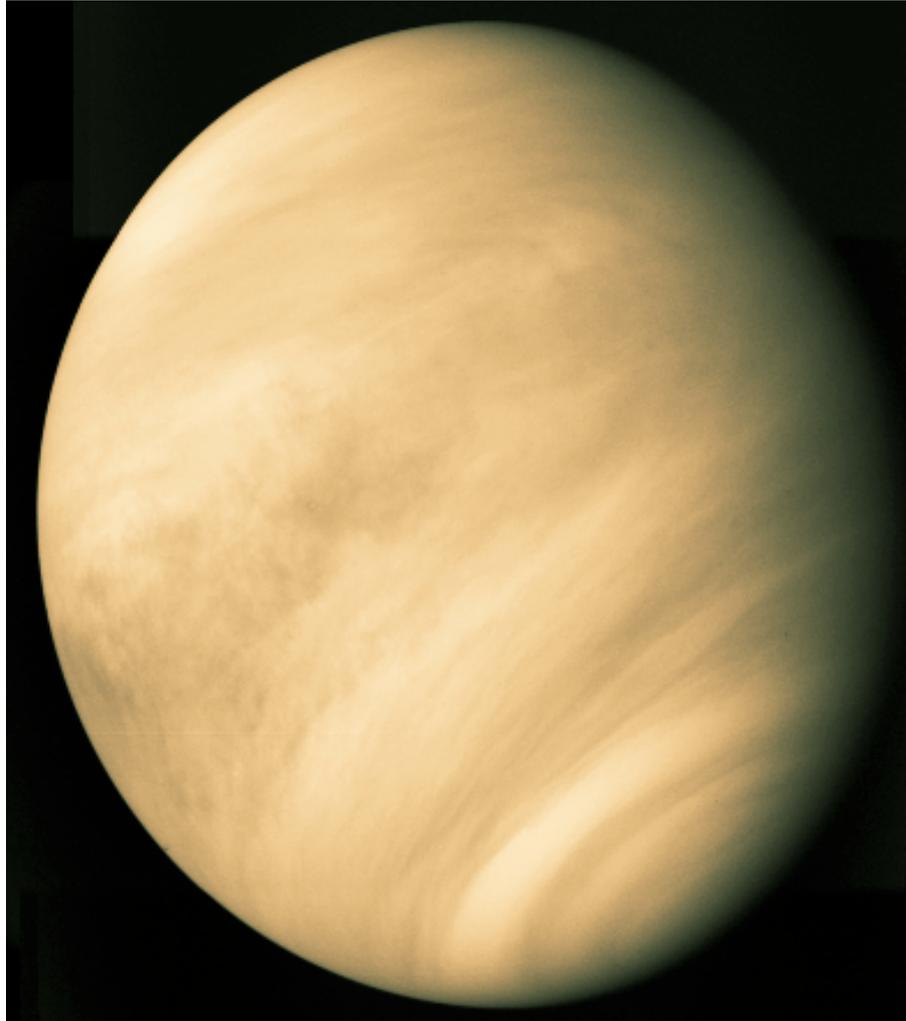
# Greenhouse Effect on Venus



Thick carbon dioxide atmosphere produces an extremely strong greenhouse effect.

Earth escapes this fate because most of its carbon and water is in rocks and oceans.

# How did Venus get so hot?



# Runaway Greenhouse Effect

Imagine moving Earth closer to the Sun, where Venus is. What would happen?

1. Temps would rise (closer to Sun)
2. Water would evaporate, leading to increased water vapor – a greenhouse gas.
3. Temps would rise more.
4. More water would evaporate, more greenhouse, temps rise more.
5. Without liquid water, carbon dioxide doesn't get washed out of the atmosphere. So carbon dioxide levels rise – another greenhouse gas. So the temps rise even more!
6. Eventually, strong UV light from the Sun would break up the water molecules in the atmosphere into hydrogen and oxygen, and the oxygen would be lost via thermal escape – no more water molecules, just carbon dioxide.

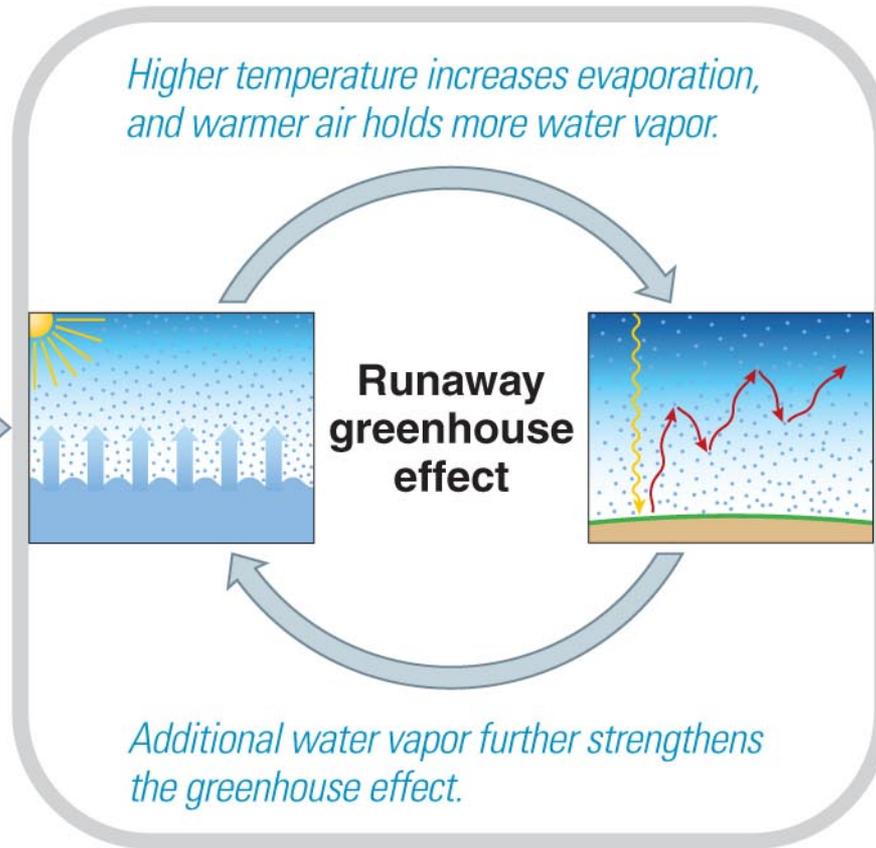
# Runaway Greenhouse Effect

**If Earth moved to Venus's orbit**

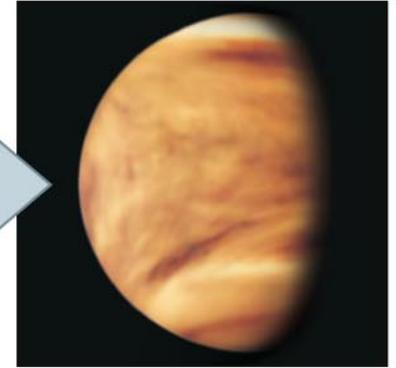
*More intense sunlight...*



*...would raise surface temperature by about 30°C.*



*Result: Oceans evaporate and carbonate rocks decompose, releasing CO<sub>2</sub>...*



*...making Earth hotter than Venus.*

- A runaway greenhouse effect would account for why Venus is so hot with so little water.

# What have we learned?

- **What is Venus like today?**
  - Venus has an extremely thick CO<sub>2</sub> atmosphere.
  - Slow rotation means little weather.
- **How did Venus get so hot?**
  - Runaway greenhouse effect made Venus too hot for liquid oceans.
  - All carbon dioxide remains in atmosphere, leading to an extreme greenhouse effect.

# 10.6 Earth's Unique Atmosphere

- Our goals for learning:
  - **How did Earth's atmosphere end up so different?**
  - **Why does Earth's climate stay relatively stable?**
  - **How is human activity changing our planet?**

# How did Earth's atmosphere end up so different?

Earth



# Four Important Questions

- Why did Earth retain most of its outgassed water?
- Why does Earth have so little atmospheric carbon dioxide, unlike Venus?
- Why does Earth's atmosphere consist mostly of nitrogen and oxygen?
- Why does Earth have an ultraviolet-absorbing stratosphere?

# Earth's Water and CO<sub>2</sub>

Earth



- Earth's temperature remained cool enough for liquid oceans to form.
- Oceans dissolve atmospheric CO<sub>2</sub>, enabling carbon to be trapped in rocks.

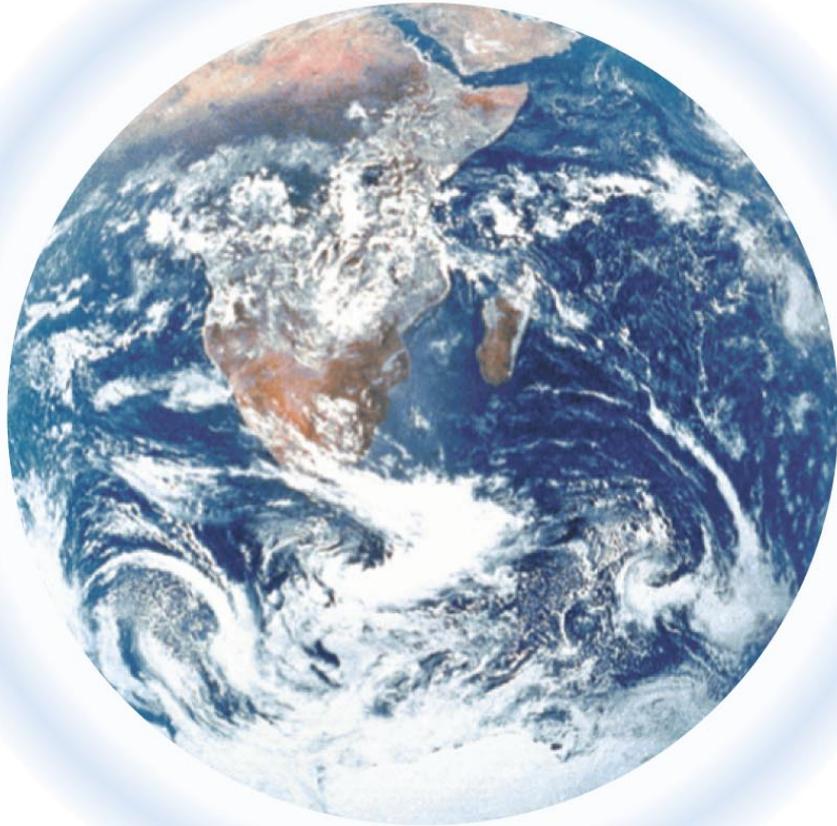
# Nitrogen and Oxygen

Earth



- Most of Earth's carbon and oxygen is in rocks, leaving a mostly nitrogen atmosphere.
- Plants release some oxygen ( $O_2$ ) from  $CO_2$  into atmosphere.

# Ozone and the Stratosphere



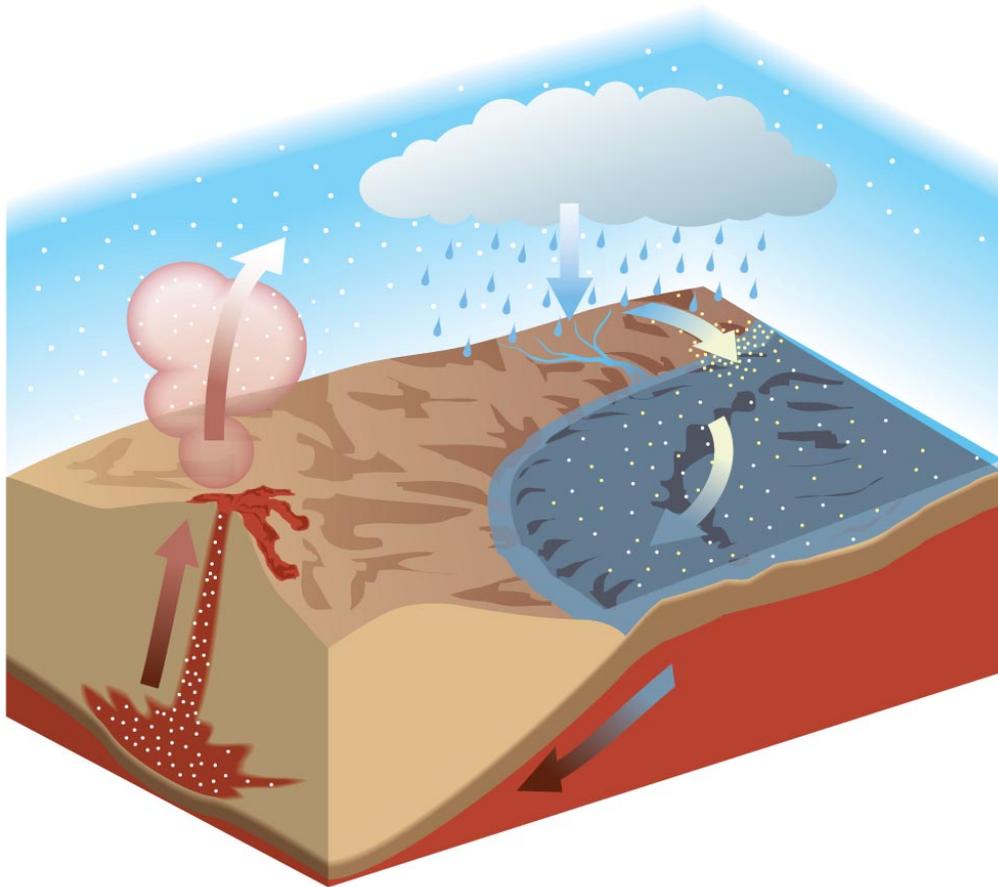
Ultraviolet light can break up  $O_2$  molecules, allowing ozone ( $O_3$ ) to form.

Without plants to release  $O_2$ , there would be no ozone in stratosphere to absorb ultraviolet light.

*Only Earth has a stratosphere!*

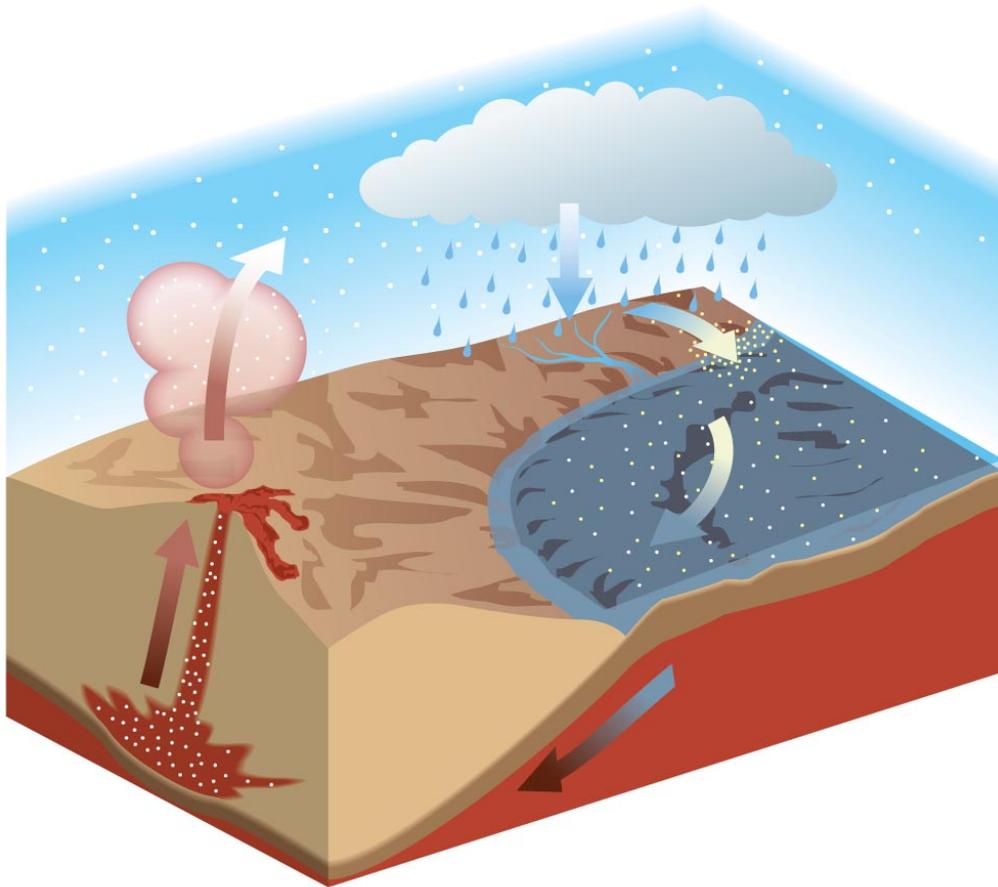


# Carbon Dioxide Cycle



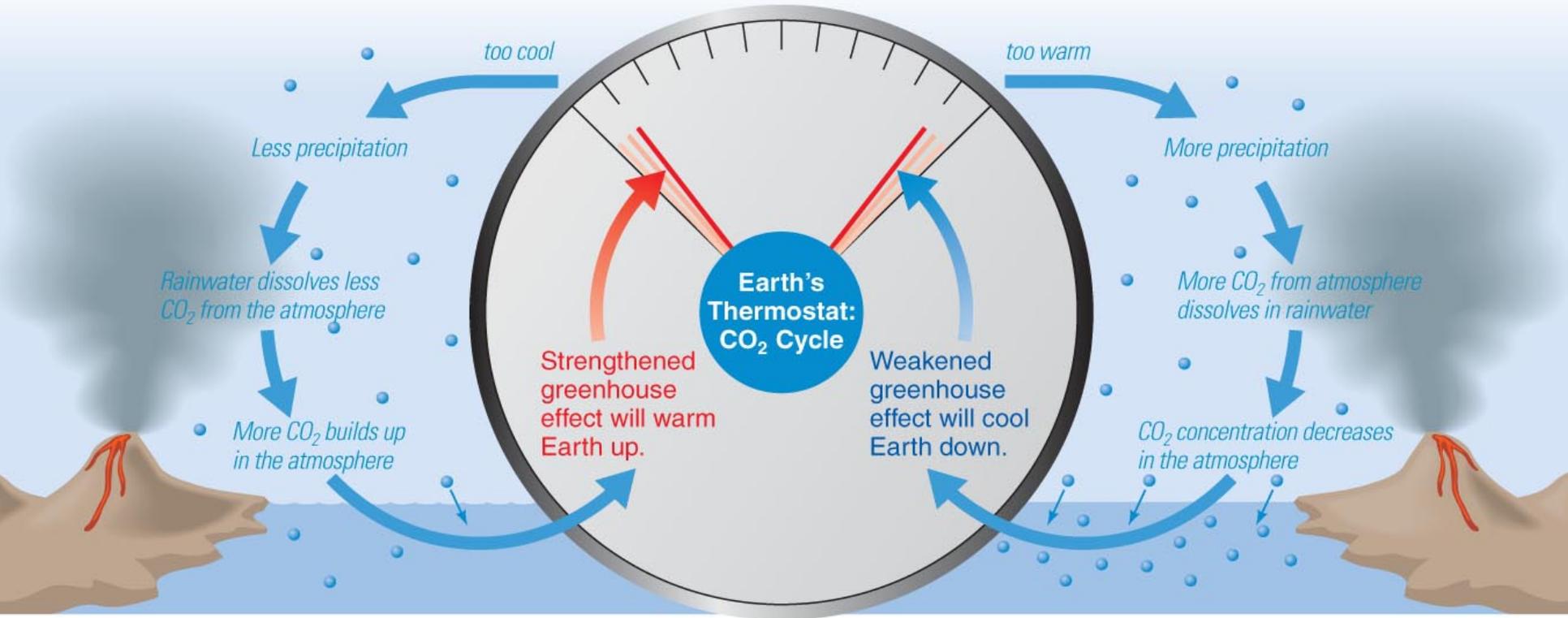
1. Atmospheric CO<sub>2</sub> dissolves in rainwater.
2. Rain erodes minerals that flow into ocean.
3. Minerals combine with carbon to make rocks on ocean floor.

# Carbon Dioxide Cycle



4. Subduction carries carbonate rock down into mantle.
5. Rock melts in mantle and  $\text{CO}_2$  is outgassed back into atmosphere through volcanoes.

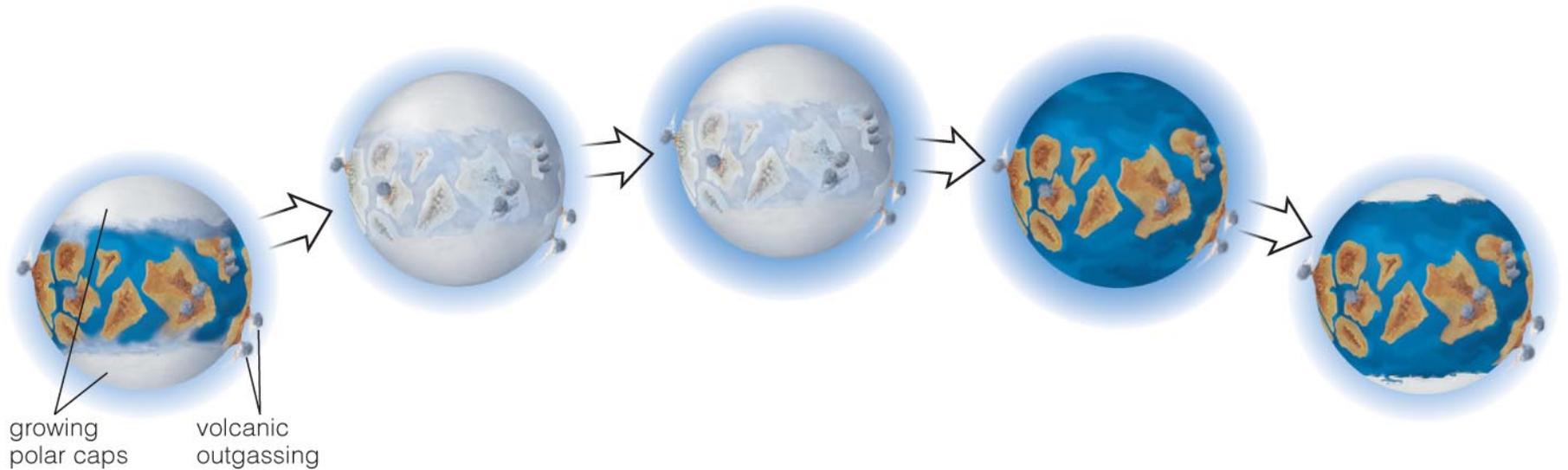
# Earth's Thermostat



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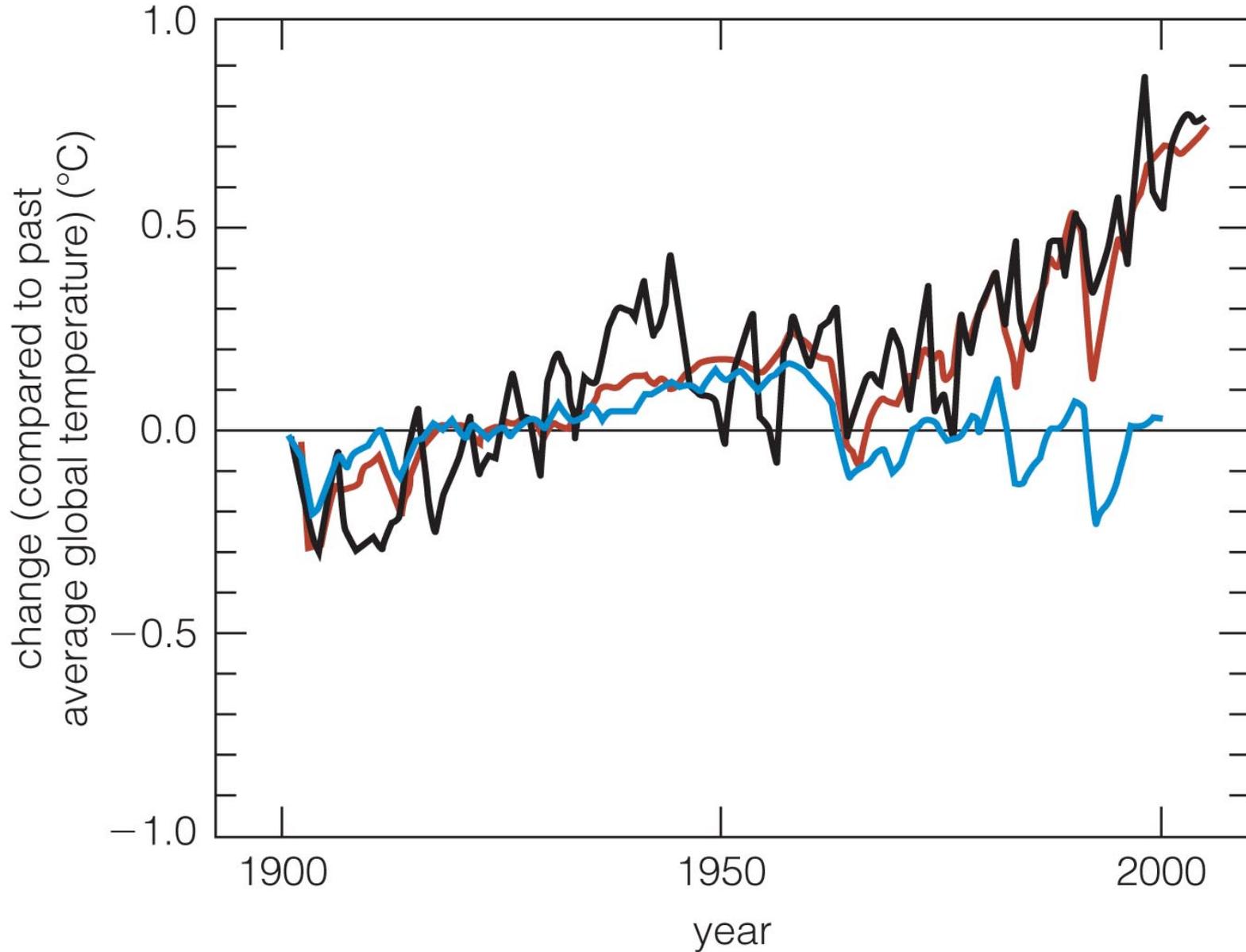
- Cooling allows CO<sub>2</sub> to build up in atmosphere.
- Heating causes rain to reduce CO<sub>2</sub> in atmosphere.

# Long-Term Climate Change



- Changes in Earth's axis tilt might lead to *ice ages*.
- Widespread ice tends to lower global temperatures by increasing Earth's reflectivity.
- CO<sub>2</sub> from outgassing will build up if oceans are frozen, ultimately raising global temperatures again.

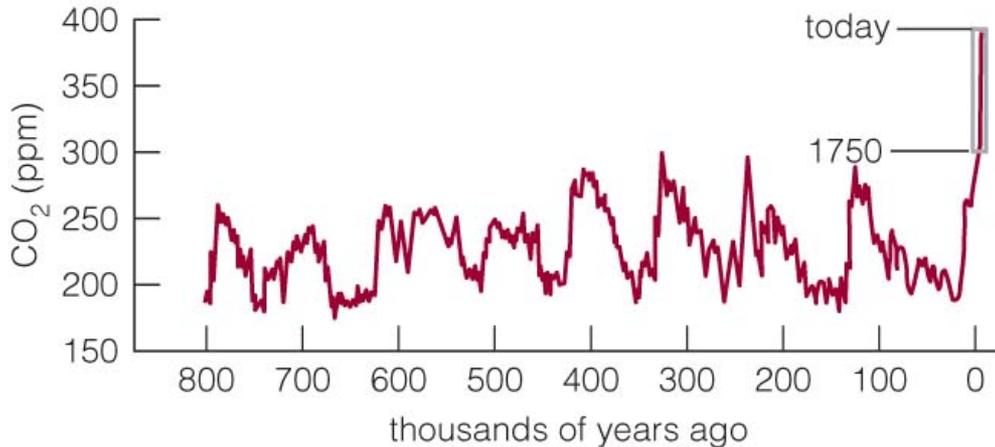
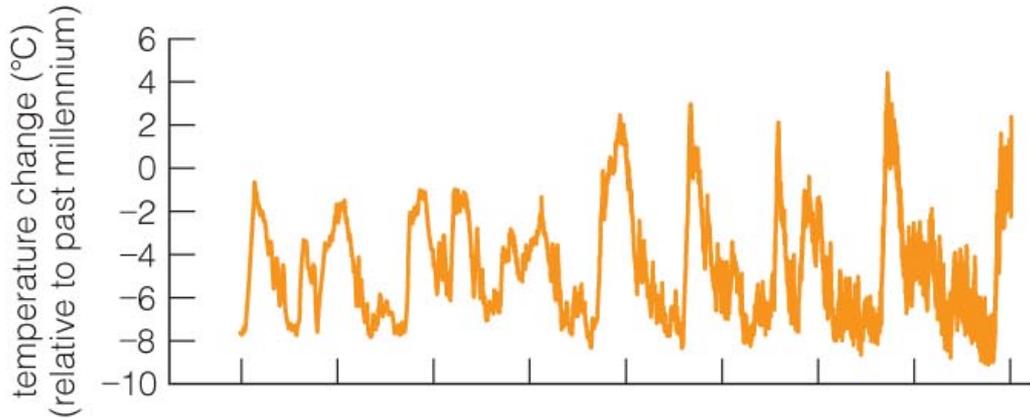
# How is human activity changing our planet?



# Global Warming

- Earth's average temperature has increased by about 1° C in past 50 years.
- The concentration of CO<sub>2</sub> is rising rapidly.
- An unchecked rise in greenhouse gases will eventually lead to global warming.

# CO<sub>2</sub> Concentration

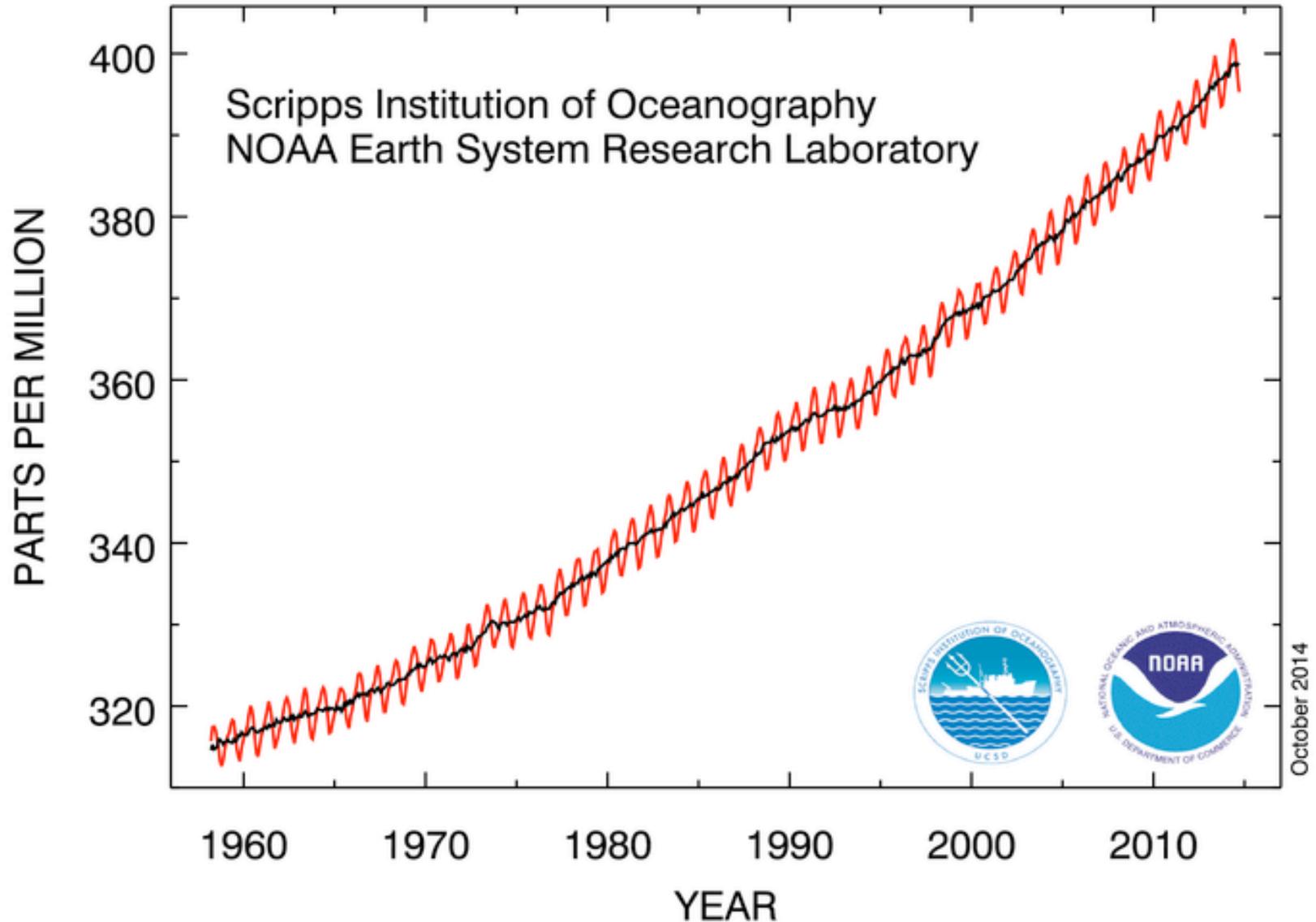


[NOAA paleo data](#)

- Global temperatures have tracked CO<sub>2</sub> concentration for last 500,000 years.
- Current CO<sub>2</sub> concentration is the highest it's been in at least 500,000 years.

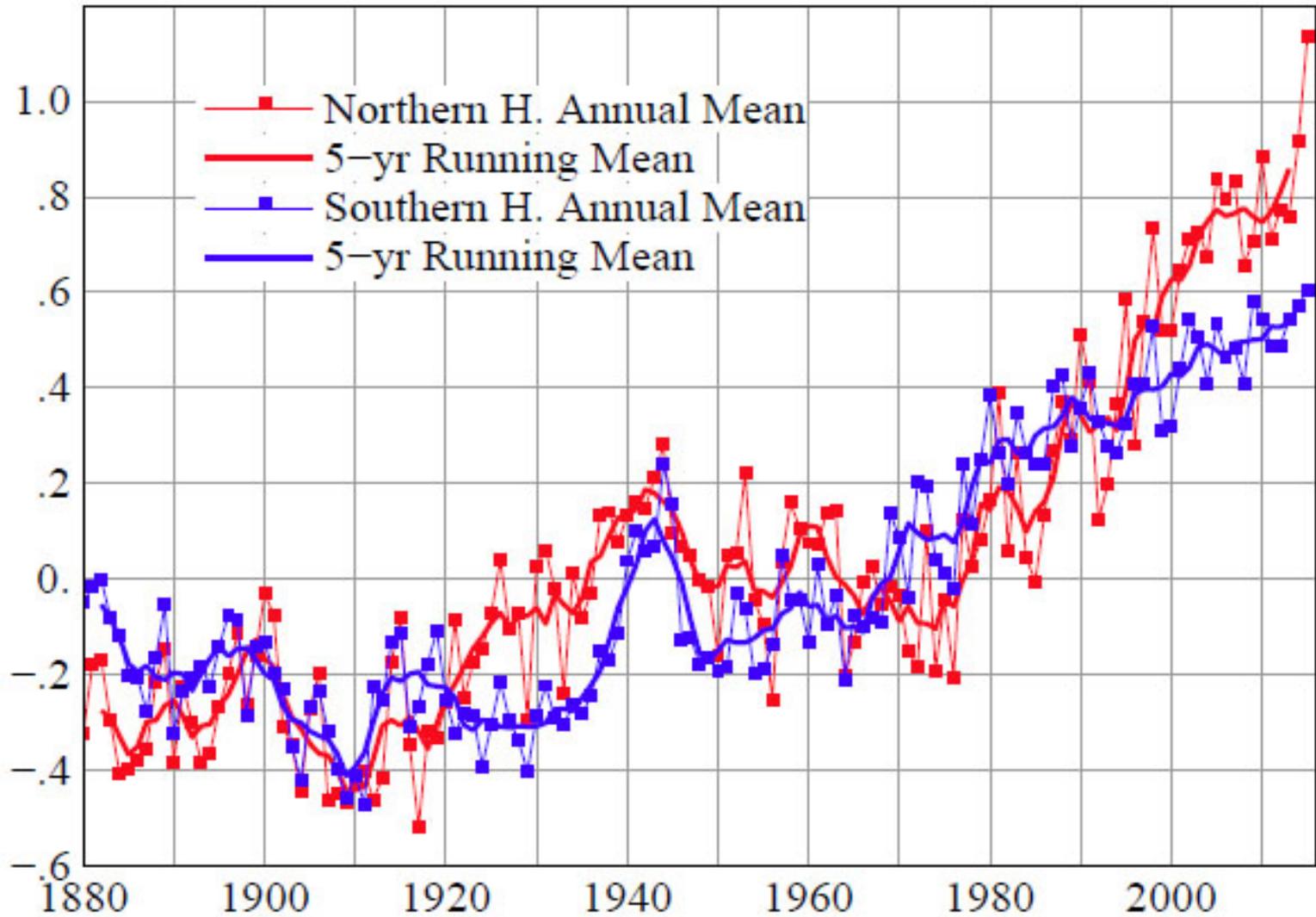
# CO<sub>2</sub> Concentration

## Atmospheric CO<sub>2</sub> at Mauna Loa Observatory



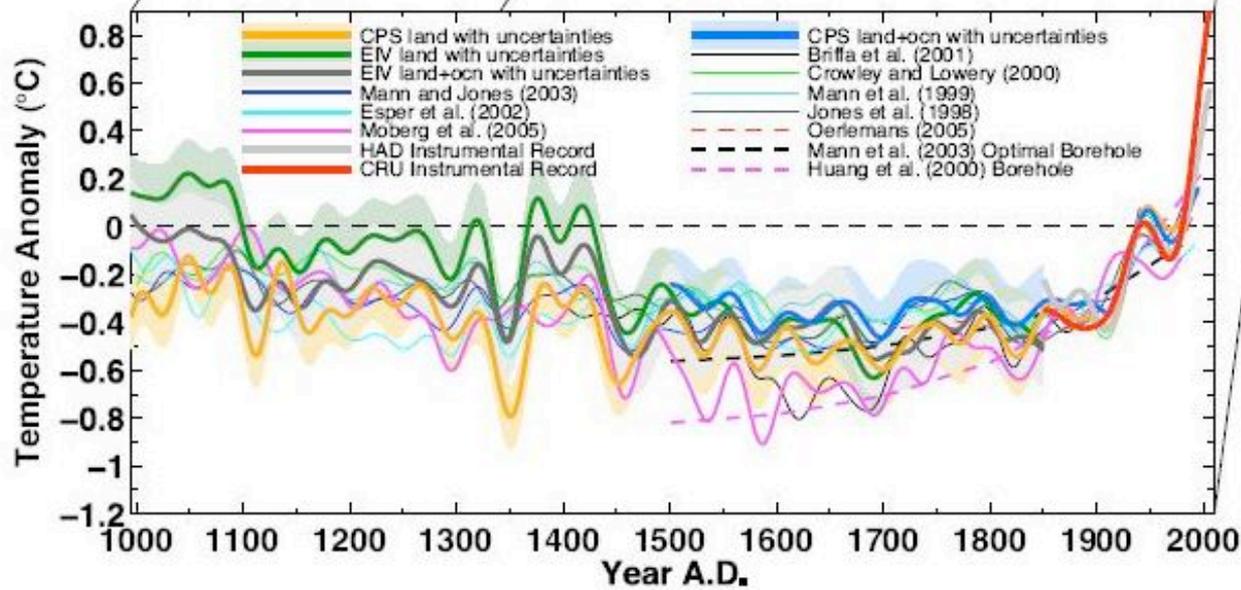
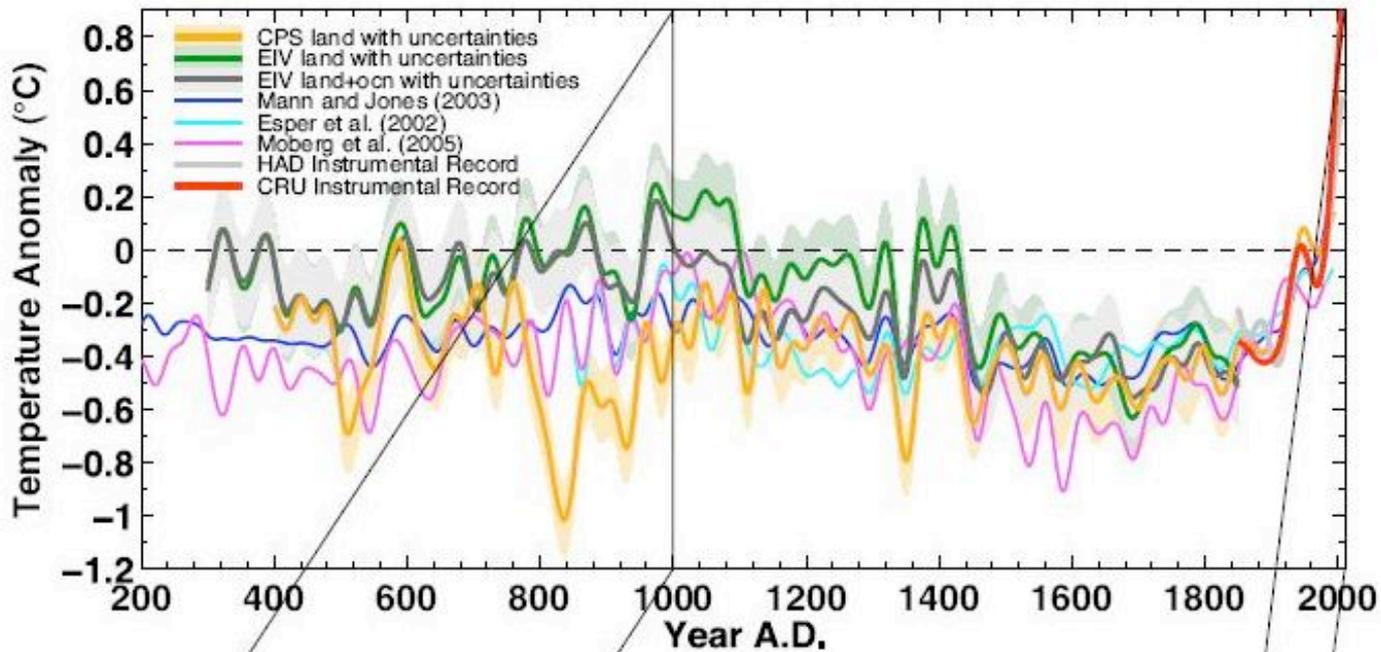
# Temperature

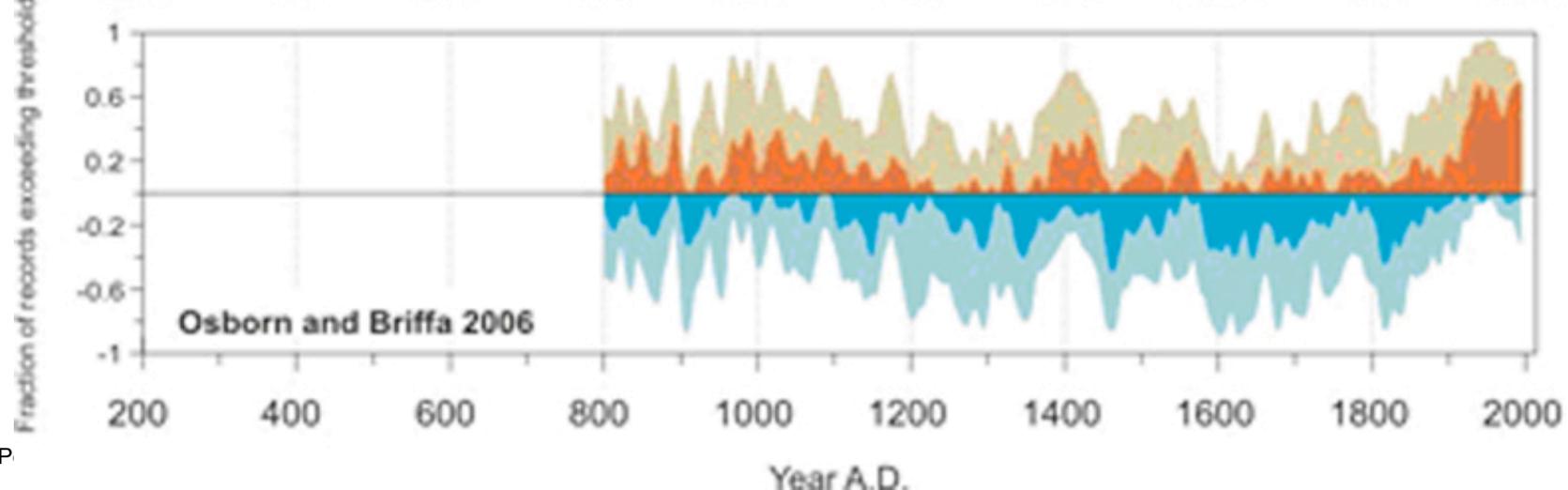
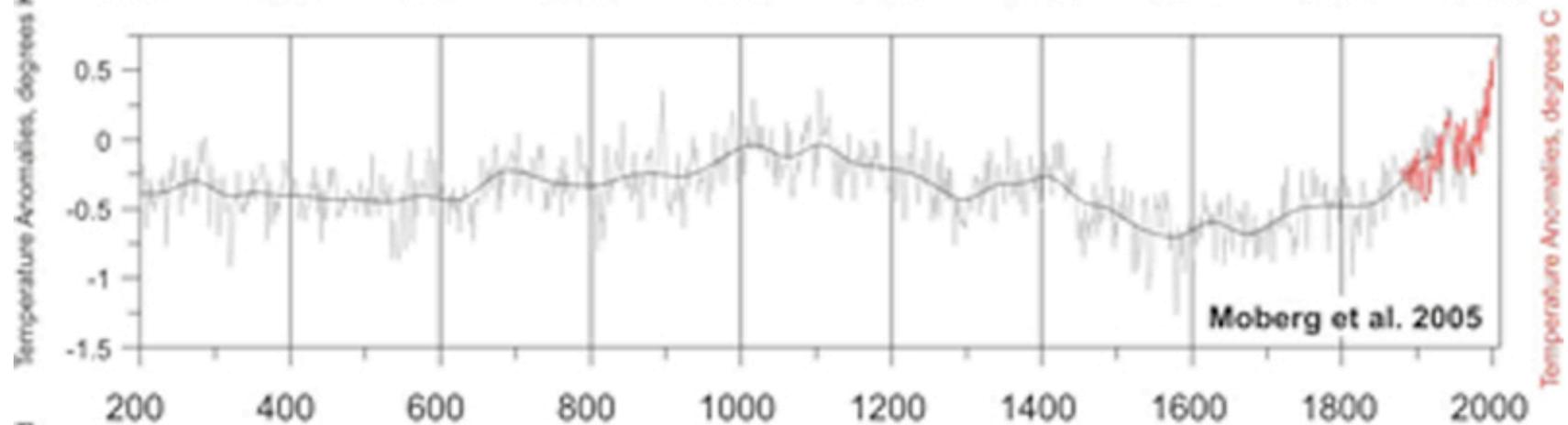
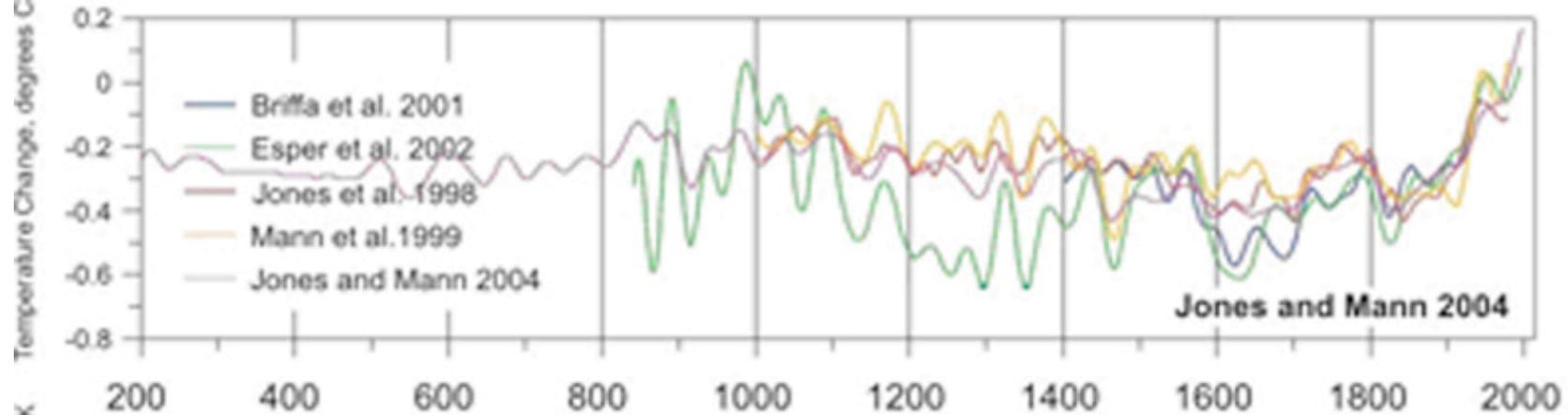
## Hemispheric Temperature Change



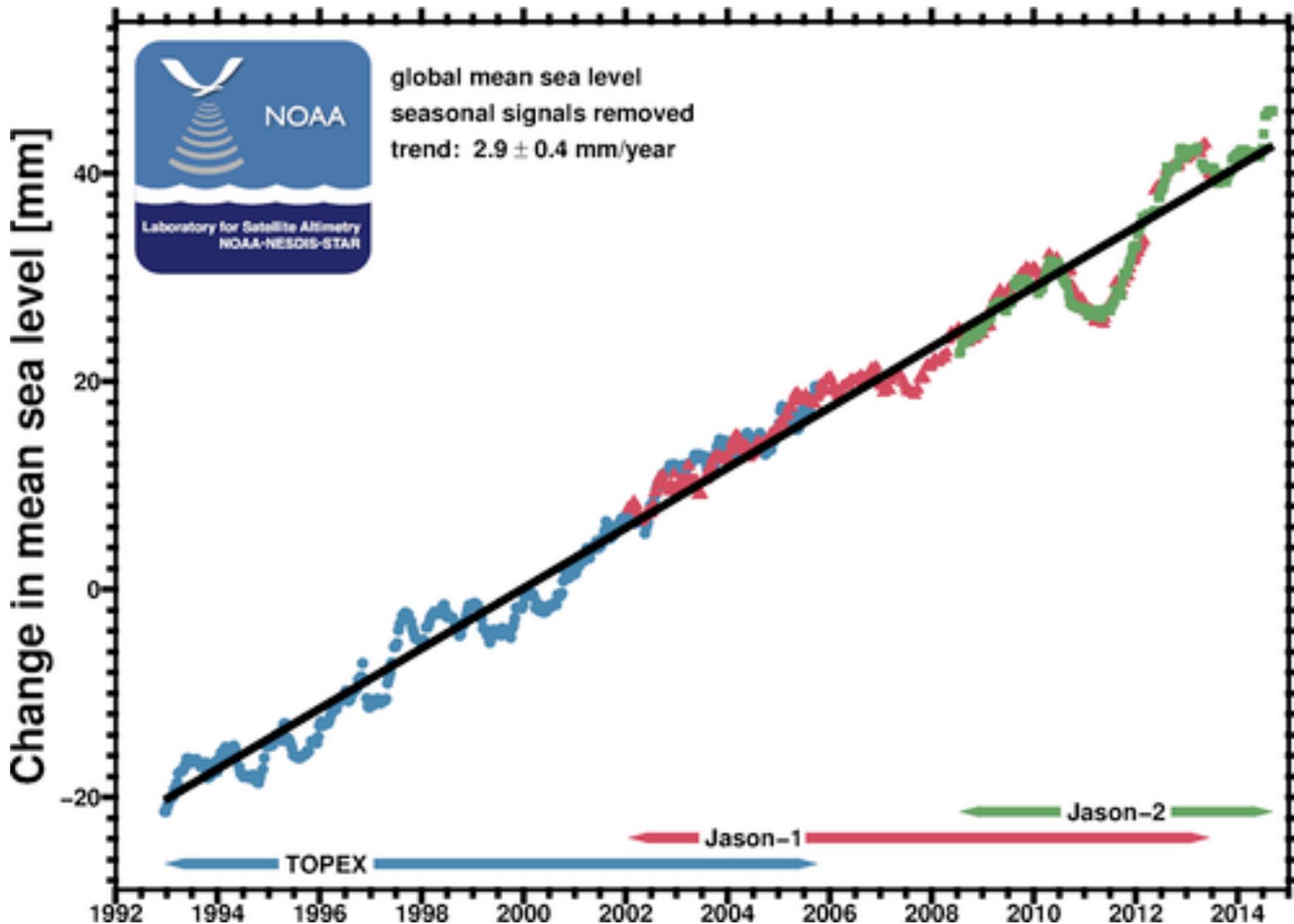
<https://data.giss.nasa.gov/gistemp/graphs/>

# Northern Hemisphere



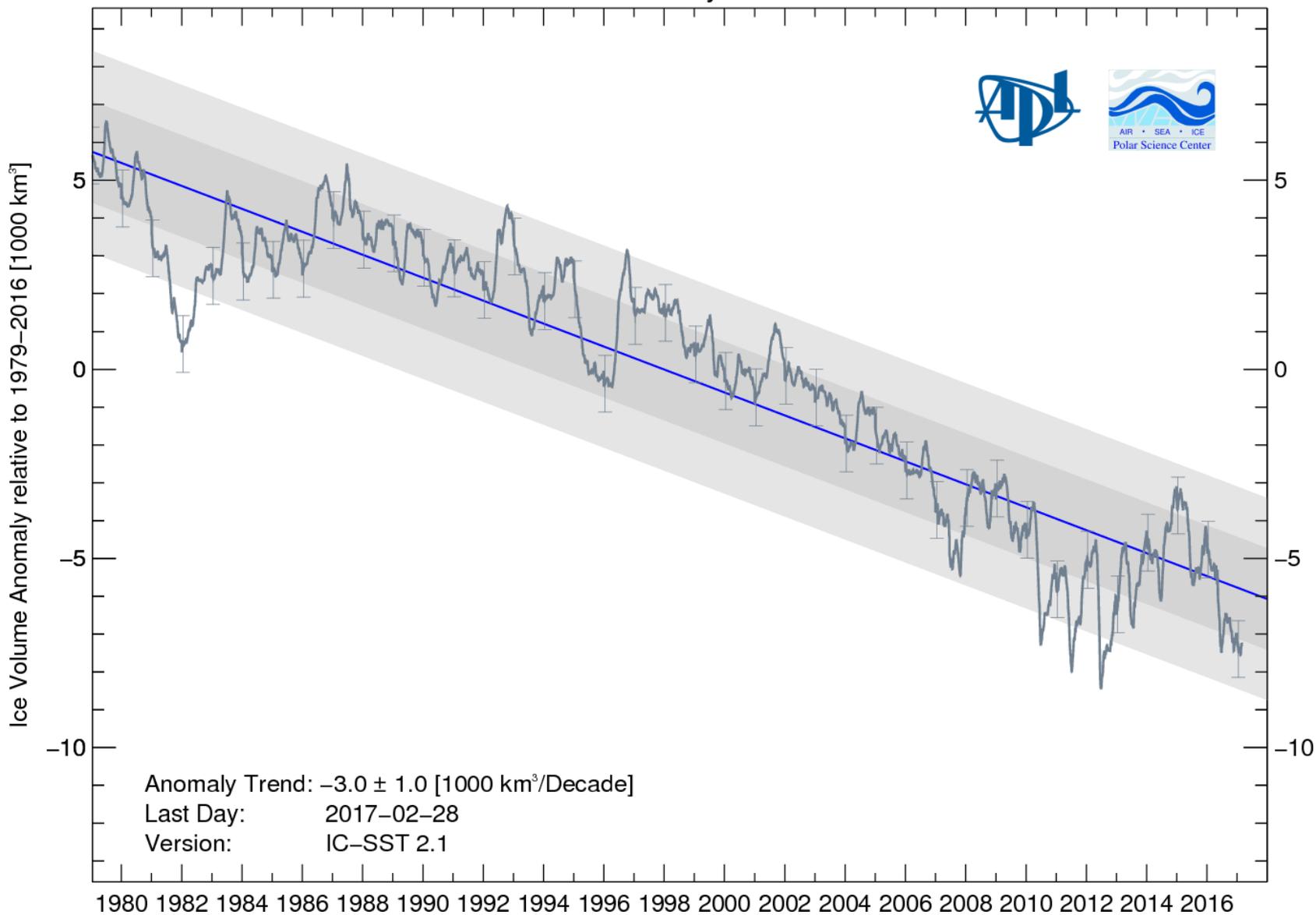


# Sea Level



# Arctic Ice

Arctic Sea Ice Volume Anomaly and Trend from PIOMAS

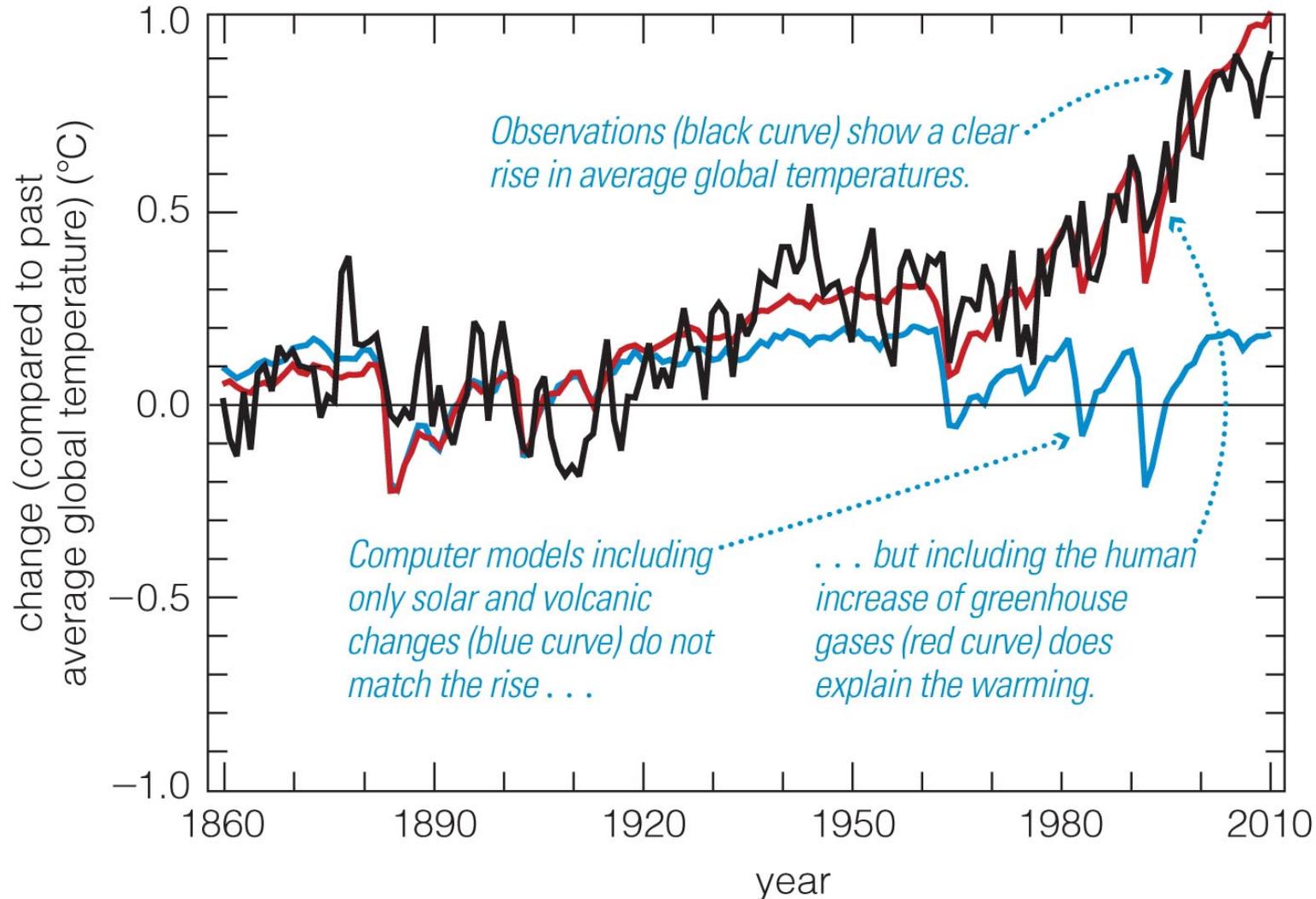


# Temperature

Global Visualization of Temperature Change

Global Modeling of Temperature Change

# Modeling of Climate Change



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- Complex models of global warming suggest that recent temperature increase is consistent with human production of greenhouse gases.

# Consequences of Global Warming

- Storms more numerous and intense
- Rising ocean levels; melting glaciers
- Local climate variations not the same as global response.
- Uncertain effects on food production, availability of fresh water

# What have we learned?

- **How did Earth's atmosphere end up so different?**
  - Temperatures are just right for oceans of water.
  - Oceans keep most CO<sub>2</sub> out of atmosphere.
  - Nitrogen remains in the atmosphere.
  - Life releases some oxygen into atmosphere.
- **Why does Earth's climate stay relatively stable?**
  - Carbon dioxide cycle acts as a thermostat.

# What have we learned?

- **How is human activity changing our planet?**
  - Destruction of ozone
  - High rate of extinction
  - Global warming from the production of greenhouse gases